# Unifying the Galilei and the Special Relativity

Alexandre Lyra \*

Universidade Federal do Rio de Janeiro, Observatório do Valongo Rio de Janeiro, RJ, 20080-090, Brazil

Marcelo Carvalho<sup>†</sup> Departamento de Matemática, Universidade Federal de Santa Catarina Florianópolis, 88.040-900, SC, Brasil and Frente Integralista Brasileira-SC (F.I.B.-S.C.) <sup>‡</sup>

#### Abstract

We present two models combining some aspects of the Galilei and the Special Relativity that leads to a unification of both relativities. This unification is founded on a reinterpretation of the absolute time of the Galilei relativity that is considered as a quantity in its own and not as mere reinterpretation of the time of the Special relativity in the limit of low velocity. In the first model, the Galilei relativity plays a prominent role in the sense that the basic kinematical laws of Special relativity, e.g. the Lorentz transformation and the velocity law, follows from the corresponding Galilei transformations for the position and velocity. This first model also provides a new way of conceiving the nature of relativistic spacetime where the Lorentz transformation is induced by the Galilei transformation through an embedding of 3-dimensional Euclidean space into hyperplanes of 4-dimensional Euclidean space. This idea provides the starting point for the development of a second model that leads to a generalization of the Lorentz transformation, which includes, as particular cases, the standard Lorentz transformation and transformations that apply to the case of superluminal frames.

# 1 Introduction

It is common to consider the Galilei relativity as the low velocity limit case of the Special Relativity (SR). This is revealed by the behavior of the Lorentz transformation when we take  $\frac{v}{c} \rightarrow 0$ , e.g.

$$\vec{x}' = \vec{x} - (1 - \gamma) \frac{\vec{x} \cdot \vec{v}}{v^2} \vec{v} - \gamma \vec{v}t \quad \longrightarrow \quad \vec{x}' = \vec{x} - \vec{v}t \tag{1}$$
$$t' = \gamma \left( t - \frac{\vec{x} \cdot \vec{v}}{c^2} \right) \quad \longrightarrow \quad t' = t \; .$$

<sup>\*</sup>e-mail: alexandr@ov.ufrj.br

<sup>&</sup>lt;sup>†</sup>Pro Brasilia fiant eximia.

<sup>&</sup>lt;sup>‡</sup>In the course of this work, M.C. wishes to register his vehement protest against those in universities across Brazil that under the false premise of vindicating better conditions and disimulating their own mediocricy sistematically go into strike, thus betraying our nation and a whole generation of brazilian students that have denied their right for education.

In this view, the absolute time of the Galilei relativity, henceforth denoted by  $\tau$ , is not considered as an independent quantity but it refers to a particular situation of SR when it is possible to identify t and t', which allows us to take  $\tau \equiv t = t'$ . It is in this sense that the absolute time is usually conceived. Therefore, the fact that one generally has  $t' \neq t$  seems to deny the possibility of having an absolute time in SR. We refer to the time of the SR as the *physical time*.

In our work we intend to reinterpret the notion of absolute time in such way that the Galilei relativity recovers its role as a theory in its own, correcting then a common view that treats the principles of the Galilei and the Special relativity as irreconcilable notions. One of the difficulties we encounter to carry out this goal is related to the way we currently understand the concept of time, which is somehow already shaped by ideas of the SR. Here, we focus our efforts on the development of two models, called Model I and II, that are built in order to allow to introduce the concept of absolute time in a general set up that also incorporates the concept of time of SR.

In model I, the fact that  $t \neq t'$  suggest us not to identify the absolute time with t but with a function depending on t and, perhaps, on the spatial location where the event occurred. Then, if for a certain instant  $\tau$  we have  $(t, \vec{x})$ ,  $(t', \vec{x}')$  representing the same event with respect to two inertial frames, we would expect the *absolute* character of  $\tau$  to imply that  $\tau(t, \vec{x}) = \tau(t', \vec{x}')$ , this expression being verified for any value of the ratio v/c. In addition, we would also expect to obtain the Lorentz transformation by replacing  $\tau(t, \vec{x})$  into the Galilei law  $\vec{x}' = \vec{x} - \vec{v}\tau$ . Having established this, the next step would be to obtain the velocity law of SR from the corresponding velocity law of the Galilei relativity,  $\frac{d\vec{x}'}{d\tau} = \frac{d\vec{x}}{d\tau} - \vec{v}$ . The fulfillment of these two laws places the Galilei relativity as a guiding principle for deriving some of the SR laws.

From a calculational perspective, the Lorentz transformation can be obtained by considering as variables the set  $\{t, \vec{x}, t', \vec{x}'\}$ , together with the equation  $\vec{x}'^2 - c^2 t'^2 = \vec{x}^2 - c^2 t^2$ . In model I, we will show that in order to combine the Galilei and the Special relativity it is sufficient to enlarge the variables' set of SR to  $\{t, \vec{x}, t', \vec{x}', \tau\}$  and to consider the pair of equations

$$\begin{cases} \vec{x}^{\,\prime 2} - c^2 t^{\,\prime 2} = \vec{x}^{\,2} - c^2 t^2 \\ \vec{x}^{\,\prime} = \vec{x} - \vec{v}\tau \;. \end{cases}$$
(2)

Here, the effect of this enlargement is to produce not only the usual Lorentz transformation but also an extra equation relating the absolute and the physical time, which gives an operational definition for the absolute time associated to the occurrence of an event.

In what concerns the structure of spacetime, we notice that both relativities are formulated in terms of a 4-dimensional space  $\mathbb{R}^4 \equiv \mathbb{R} \times \mathbb{R}^3$ . In the Galilei view, two observers belonging to inertial frames S, S' describe the occurence of an event respectively as a point  $(\tau, \vec{x}) \in \mathbb{R}_S \times \mathbb{R}^3_S \equiv$  $\mathbb{R}^4_S$ , and  $(\tau, \vec{x}') \in \mathbb{R}_{S'} \times \mathbb{R}^3_{S'} \equiv \mathbb{R}^4_{S'}$ . Here, due to the nature of the absolute time, for a fixed  $\tau$ , the Galilei transformation reduces itself to a map  $\mathbb{R}^3_S \to \mathbb{R}^3_{S'}$ .

In the SR view, the observers describe the event as a point  $(t, \vec{x}) \in \mathbb{R}_S \times \mathbb{R}_S^3$ , and  $(t', \vec{x}') \in \mathbb{R}_{S'} \times \mathbb{R}_{S'}^3$ , where now, due to the nature of the physical time, the Lorentz transformation is not restricted to be a map  $\mathbb{R}_S^3 \to \mathbb{R}_{S'}^3$ , as before, but it is considered as a map on the whole

space,  $\mathbb{R}_{S}^{4} \to \mathbb{R}_{S'}^{4}$ . In model I, the existence of the absolute time allows us to unveil an additional structure present in the SR spacetime that is already present in the Galilei view. In fact, when we express the absolute time in terms of the physical time,  $\tau(t)$ ,  $\tau(t')$ , we will show that we can define embeddings  $\mathbb{R}_{S}^{3} \stackrel{i_{\tau}}{\to} \mathbb{R}_{S'}^{4}$ ,  $\mathbb{R}_{S'}^{3} \stackrel{i'_{\tau}}{\to} \mathbb{R}_{S'}^{4}$  in terms of certain hyperplanes  $\sigma_{\tau} \subset \mathbb{R}_{S}^{4}$ ,  $\sigma'_{\tau} \subset \mathbb{R}_{S'}^{4}$ . The Lorentz transformation is then seen as a map  $\sigma_{\tau} \to \sigma'_{\tau}$  between these 3-dimensional hyperplanes that is induced by the Galilei transformation acting on  $\mathbb{R}_{S}^{3} \to \mathbb{R}_{S'}^{3}$ , for a certain  $\tau$ . In addition, the spacetime splits as the union of these hyperplanes, e.g.  $\mathbb{R}_{S}^{4} = \cup_{\tau} \sigma_{\tau}$ . In the spacetime of Galilei relativity this splitting is self-evident and corresponds to the hyperplanes of  $\tau = \text{constant}$ .

Inspired by the equation that defines the embeddings  $\mathbb{R}^3_S \stackrel{i_{\tau}}{\hookrightarrow} \mathbb{R}^4_S$ ,  $\mathbb{R}^3_{S'} \stackrel{i'_{\tau}}{\hookrightarrow} \mathbb{R}^4_{S'}$ , we develop Model II by taking  $\alpha x^0 - \vec{\beta} \cdot \vec{x} = \alpha x'^0 + \vec{\beta} \cdot \vec{x}'$  as our fundamental equation, where  $\alpha, \vec{\beta}$  are arbitrary parameters that will allow us to define a generalization of the Lorentz transformation. In model II the absolute time is introduced by modifying the previous equation to

$$\tau \equiv \alpha x^0 - \vec{\beta} \cdot \vec{x} = \alpha x'^0 + \vec{\beta} \cdot \vec{x}'$$

Here, the Galilei relativity arises assuming in addition the relation  $\vec{x}' = \vec{x} - \vec{v}\tau$ . This is sufficient to reduce the generalized transformation to the particular form of the standard Lorentz transformation, showing then the consistency of both relativities. In Model II, however, we will show that we are allowed to have the Lorentz transformation weather or not we consider the Galilei relativity, a situation that doesn't happen in model I.

Some work [1, 2] deal with the Galilei and the Special relativity but focus on opposite goals not proposing a scheme for unifying both relativities. In fact, in [1] the assumption that the Galilei law of velocity applies to the motion of bodies, signals and forces leads to inconsistencies that are solved by the introduction of the postulate of the constancy of the speed of light, which ultimately leads to SR. In [2], it is established a transformation relating the Galilei and the Minkowski-Einstein coordinates, which may signalize that relativistic effects are due to motion relative to an actual 3-space <sup>1</sup>. In our work we follow another direction as our goal is to harmonize both relativities by a convenient treatment of the absolute time and the fundamentals laws of both relativities. We are not aware of any work devoted to this issue, therefore, we hope our work may provide a convenient starting point for further investigation on this topic.

Our work is organized as follows. The Model I is developed in section 3. In section 3.1 we introduce a set of assumptions settling the main properties of the physical space and time that are necessary to develop a model incorporating both the Galilei and the Special Relativity. We axiomatize the existence of two times, the *absolute* time of the Galilei relativity and the *physical* time of SR, and we discuss an important issue, due to C. Moller, concerning the correct interpretation of the physical space where equation (1) is defined. In section 3.2 we obtain the Lorentz transformation from the Galilei law  $\vec{x}' = \vec{x} - \vec{v}\tau$  and the assumptions of section 3.1. In section 3.4 we obtain from the law of velocity transformation of Galilei relativity the corresponding velocity law of SR. We also discuss a problem associated to the so-called Thomas

<sup>&</sup>lt;sup>1</sup>The author continues this development in a model that is known by *Process Physics*.

precession. In section **3.5** we give a geometric interpretation for the spacetime of SR, and show how the concept of absolute time allows to define a partition of the spacetime in terms of certain 3-dimensional hyperplanes. The Model II is developed in section **4.** In section **4.1** we establish a set of assumptions that is slightly different than the assumptions of Model I, namely in what concerns the introduction of the Special and the Galilei relativity (respectively numbered as assumptions III and IV). In section **4.2** we derive a transformation that we call Generalized Lorentz Transformation, and in section **4.3** we analyze the velocity transformation associated to it. We then show that the Generalized Lorentz Transformation includes, as particular cases, the standard Lorentz transformation together with transformations that may be used when we have superluminal frames. We then analyze in section **4.4** how to introduce the Galilei relativity in the framework of Model II. In section **4.5** we analyze the conservation of momentum and obtain the corresponding relativistic expression for the mass. Finally, in section **5** we analyze how elementary considerations from the differential structure of the projective space  $\mathbb{RP}(5)$  allows us to think on the concept of the absolute time as associated to one of the dimensions of Euclidean space  $\mathbb{R}^4$ .

# 2 Basic Definitions

We recall some basic definitions that are necessary for the statement of our assumptions.

An event is any physical occurrence taking place on a certain location and on a certain instant. We adopt the standard definitions of observer and reference frame as stated concisely in [3], i.e. by an observer we understand any entity equipped with a standard rod and a standard clock that allows for the measurement of length and time (in fact, the physical time as we will introduce in the assumption II below) and that is able to communicate with other observers by means of light signals. A reference frame is understood as an infinite set of observers each one at rest relative to the other and having their standard clocks synchronized. Abstractly, the infinite set of observers composing a reference frame is idealized in such way that for any event there is an observer present on the same location where the event took place, which then establishes the space coordinate of the event. This observer also determines, by the reading of his clock, the instant of time when the event occurred. We will consider reference frames endowed with a rectangular coordinate system. By an *inertial frame* we understand any frame in which a body free of forces is unaccelerated.

We use the notation  $(t_{SP}, \vec{x}_{SP})$  to denote the description of an event P relative to a frame S. We write  $(t_{SP}, \vec{x}_{SP}) \sim (t_{S'P}, \vec{x}_{S'P})$  to indicate the same event P as seen by the frames S, S'. In the particular case when one frame S analyzes the movement of another frame S', we will assume the frame S' is entirely represented by its origin, and write  $\vec{x}_{SS'}$  for the position of S' relative to S, and  $t_{SS'}$  for the time measured by S. Finally, in our work by a Lorentz transformation we always mean a Lorentz boost, except in section 5 where a Lorentz boost attains its precise meaning as a particular Lorentz transformation.

# 3 Model I

### 3.1 The Assumptions

In what concerns *space* and *time* our basic assumptions are as follows.

### I. Space

Each inertial frame describes space as being an euclidean 3-dimensional vector space.

#### II. Time

We model time as any variable that can be used to describe what we intuitively understand as the "instant when events occur". We distinguish two kind of choices as follows. By *physical time*, denoted by t, we understand a choice for the time variable that can be measured with the use of standard devices, like atomic clocks, under suitable arrangements (e.g. when they are the clocks of a reference frame and are all conveniently synchronized etc.). By *absolute time*, denoted by  $\tau$ , we understand a choice for the time variable having the property that given an event all observers assign to it the same value for the instant when the event occurred <sup>2</sup>. As a principle, we assume that any variable that serves to describe time may be expressed in terms of the physical time in such way that by the measurement of the latter one can determine the value of the former. Therefore, for any frame in which one knows how to shift from  $\tau(t) \leftrightarrow t(\tau)$ it is possible to describe an event writing its coordinates as  $(t, \vec{x})$  or  $(\tau, \vec{x})$ .

## III.

Given two inertial frames S, S' moving with relative velocity  $\vec{v}$ , and an event P whose coordinates are  $(t_{SP}, \vec{x}_{SP}), (t_{S'P}, \vec{x}_{S'P})$  relative to S, S', we have

$$\vec{x}_{SP}^2 - c^2 t_{SP}^2 = \vec{x}_{S'P}^2 - c^2 t_{S'P}^2 .$$
(3)

When we consider speed calculated in terms of derivatives relative to the physical time, the relation (3) expresses the constancy of the speed of light.

### IV. The Galilean Relativity Principle

Given two inertial frames S, S' moving with relative velocity  $\vec{v}$ , and an event P whose coordinates are  $(\tau, \vec{x}_{SP}), (\tau, \vec{x}_{S'P})$  relative to S, S', we have *componentwise* that

$$\vec{x}_{S'P} = \vec{x}_{SP} - \vec{v}_{SS'}\tau \ . \tag{4}$$

**V.** The relation between  $(t_{SP}, \vec{x}_{SP})$  and  $(t_{S'P}, \vec{x}_{S'P})$  is linear.

#### **Remarks:**

(i) In assumption I, it is assumed that observers in inertial frames see physical space endowed with an euclidean structure. This assumption is supported by the lack of evidence signalyzing deviations of space properties from the euclidean structure (see the discussion of A. P. French in [4], pp. 59-61). The position of an event P relative to inertial frames S, S' is written as

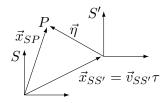
 $<sup>^{2}</sup>$ Later, through the analysis of their transformation properties, we will identify the physical time as the ordinary time of the special relativity, while the absolute time will be identified with the time of the Galilei relativity.

 $\vec{x}_{SP} \equiv (x^1, x^2, x^3) \in \mathbb{R}^3_S, \ \vec{x}_{S'P} \equiv (x'^1, x'^2, x'^3) \in \mathbb{R}^3_{S'}$  and since each frame is endowed with its own coordinate system the spaces  $\mathbb{R}^3_S$  and  $\mathbb{R}^3_{S'}$  are conceived as *distinct* spaces, therefore, we must understand the relation between the vectors  $\vec{x}_{SP}$  and  $\vec{x}_{S'P}$  established in (4) not as a vector equation defined in one and the same vector space. However, it is possible to do so if we reinterpret the vectors  $\vec{x}_{SP}$  and  $\vec{x}_{S'P}$  as follows <sup>3</sup>. Given two frames S, S' we assume there is a *single* space  $\mathbb{R}^3_{SS'}$  where the components  $(x^1, x^2, x^3), (x'^1, x'^2, x'^3)$  of the vectors  $\vec{x}_{SP}, \vec{x}_{S'P}$  are seen as images of maps

$$\mathbb{R}^{3}_{S} \to \mathbb{R}^{3}_{SS'} : \vec{x}_{SP} \to (x^{1}, x^{2}, x^{3})$$
$$\mathbb{R}^{3}_{S'} \to \mathbb{R}^{3}_{SS'} : \vec{x}_{S'P} \to (x'^{1}, x'^{2}, x'^{3})$$

and such that for a fixed  $\tau$  they satisfy equation (4), i.e. we can think of the vectors  $\vec{x}_{SP}$ ,  $\vec{x}_{S'P}$  as vectors in  $\mathbb{R}^3_{SS'}$ , which justify using the same notation for them. Equation (4) is now understood as an equation defined in  $\mathbb{R}^3_{SS'}$ .

From the perspective of the Galilei relativity, it is clear that the vector  $\vec{x}_{S'P} \in \mathbb{R}^3_{S'}$  can be measured directly from the observers of the frame S that at the instant  $\tau$  are placed respectively at the positions corresponding to the origin of the frame S' and the event P, i.e. componentwise we identify  $\mathbb{R}^3_{S'} \ni \vec{x}_{S'P} \leftrightarrow \vec{\eta} \in \mathbb{R}^3_S$ ,



where

$$\vec{\eta} := \vec{x}_{SP} - \vec{x}_{SS'} = \vec{x}_{SP} - \vec{v}_{SS'} \tau$$
.

This corresponds to the same equation (4), now referred entirely to the space  $\mathbb{R}^3_S$ . This procedure is equivalent to identify  $\mathbb{R}^3_{SS'}$  with  $\mathbb{R}^3_S$ . A similar procedure allows to identify  $\mathbb{R}^3_{SS'}$  and  $\mathbb{R}^3_{S'}$ . The assumption that spaces  $\mathbb{R}^3_S$ ,  $\mathbb{R}^3_S$ , are euclidean is then consistent with the interpretation we obtain for the equation (4) through the identification  $\mathbb{R}^3_S \simeq \mathbb{R}^3_{SS'}$ ,  $\mathbb{R}^3_{S'} \simeq \mathbb{R}^3_{SS'}$ .

A different picture emerges in SR. Instead of (4) we have the relation <sup>4</sup>

$$\vec{x}_{S'P} = \vec{x}_{SP} - (1 - \gamma) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} \vec{v}_{SS'} - \gamma \vec{v}_{SS'} t_{SP} .$$
(5)

Here, the vectors  $\vec{x}_{SP} \in \mathbb{R}^3_S$ ,  $\vec{x}_{S'P} \in \mathbb{R}^3_{S'}$  are mapped into an *abstract* space  $V_{SS'}$  where we understand the equation (5) is defined. From the perspective of SR and according to the frame

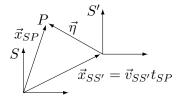
<sup>&</sup>lt;sup>3</sup>This reasoning is originally due to Moller [5].

<sup>&</sup>lt;sup>4</sup>It seems it was C. Moller (see section 2.4 in [5]) who first noticed the importance of giving a correct interpretation for equation (5) as an equation established in an abstract vector space, where it would make sense to relate the vectors  $\vec{x}_{SP}$  and  $\vec{x}_{S'P}$  that, in principle, belong to different spaces. C. Moller did so thinking exclusively from the perspective of the SR. However, the same question is already manifest from the perspective of the Galilei relativity if we intend to interpret correctly equation (4).

S, if the the event P occurs at the instant  $t_{SP}$  the observers on the frame S that occupy respectively the positions of the origin of the frame S' and the event P allow us to define a vector  $\vec{\eta}$ 

$$\vec{\eta} := \vec{x}_{SP} - \vec{v}_{SS'} t_{SP} , \qquad (6)$$

which is not componentwise equal to the vector  $\vec{x}_{S'P}$  as measured by the frame S'.



In fact, using the relation

$$t_{S'P} = \gamma \left( t_{SP} - \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{c^2} \right)$$

together with (5) to write  $\vec{x}_{SP}$  in terms of  $\vec{x}_{S'P}, t_{S'P}$  we obtain componentwise that

$$\vec{\eta} = \vec{x}_{S'P} + \frac{1-\gamma}{\gamma} \frac{\vec{x}_{S'P} \cdot \vec{v}_{S'S}}{v_{S'S}^2} \vec{v}_{S'S}$$

therefore, from the perspective of SR, we cannot identify the vectors  $\vec{\eta}$  and  $\vec{x}_{S'P}$  as we did before. Physically, this means that measurements performed by the observers of the frame S are not sufficient to identify the equation (6) with the equation (5) of  $V_{SS'}$ . We then distinguish the spaces  $\mathbb{R}^3_{SS'}$  and  $V_{SS'}$  in the sense that we cannot identify  $\mathbb{R}^3_S$  or  $\mathbb{R}^3_{S'}$  with  $V_{SS'}$  as we did in the case of the Galilei relativity.

(ii) In assumption III we consider equation (3) holds for *every* event, not only for the events associated to the movement of a light ray. As we know [6], in the standard treatment of SR if we assume that (3) is verified for a light ray then from the assumption of the linearity of the transformation involving  $\vec{x}, t$  and  $\vec{x}', t'$  we obtain that condition (3) is also verified for any event. In our approach, since we work essentially with two equations (3, 4) and an extended set of variables  $\{\tau, t, t', \vec{x}, \vec{x}'\}$  there is no guarantee that assuming (3) is true for a light ray would imply its validity for all events <sup>5</sup>. Therefore, in the development of our model it is necessary to assume from the beginning that (3) is verified for *all* events and not only for those events associated to the movement of a light ray. This will become evident in the derivation shown in section 3.2.

(iii) The physical and the absolute time introduced in assumption II assume their specific characteristic from the conditions they have to obey in assumptions III and IV. These conditions identify t as the time of Special Relativity and  $\tau$  as the time of Galilei relativity.

(iv) Whenever we set a transformation between two frames we assume that at  $\tau = 0$ , or equivalently at t = t' = 0, the origins of both frames coincide and their coordinate axis are parallel.

<sup>&</sup>lt;sup>5</sup>For the case of the standard treatment of the SR, Landau gives an heuristic argument justifying it in [7], pp.5, while Einstein put it axiomatically in [8], appendix one.

(v) In assumption V, by a linear relation between the pairs  $(t, \vec{x}), (t', \vec{x}')$  we mean that t' and  $\vec{x}'$  depend only on the first order power of  $t, \vec{x}$ .

#### **3.2** Deriving the transformation

In order to obtain the transformations we start from assumption IV

$$\vec{x}_{S'P} = \vec{x}_{SP} - \vec{v}_{SS'}\tau\tag{7}$$

which gives

$$\vec{x}_{S'P}^2 = \vec{x}_{SP}^2 - 2\vec{x}_{SP} \cdot \vec{v}_{SS'}\tau + v_{SS'}^2\tau^2 .$$

From assumption III we obtain

$$v_{SS'}^2 \tau^2 - 2\vec{x}_{SP} \cdot \vec{v}_{SS'} \tau + c^2 (t_{SP}^2 - t_{S'P}^2) = 0$$

that allows us to write

$$\tau = \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'} \pm \sqrt{(\vec{x}_{SP} \cdot \vec{v}_{SS'})^2 - v_{SS'}^2 c^2 (t_{SP}^2 - t_{S'P}^2)}}{v_{SS'}^2} \,. \tag{8}$$

Now, according to assumption V, we look for a relation between  $t_{SP}$  and  $t_{S'P}$  such that the corresponding expression for  $\tau$  given in (8), when replaced into (7), results in a transformation involving at most terms to the first power in  $\vec{x}_{S'P}$ ,  $\vec{x}_{SP}$ . In order to obtain that, let us consider the following particular relation between  $t_{SP}$  and  $t_{S'P}$ , e.g.

$$t_{S'P} = a t_{SP} + b \vec{x}_{SP} \cdot \vec{v}_{SS'} \tag{9}$$

which gives  $t_{SP}^2 - t_{S'P}^2 = (1 - a^2)t_{SP}^2 - 2ab\,\vec{x}_{SP}\cdot\vec{v}_{SS'}t_{SP} - b^2(\vec{x}_{SP}\cdot\vec{v}_{SS'})^2$ . The constants a and b must be chosen in such way that  $(\vec{x}_{SP}\cdot\vec{v}_{SS'})^2 - v_{SS'}^2c^2(t_{SP}^2 - t_{S'P}^2)$  becomes a perfect square, i.e.

$$(\vec{x}_{SP} \cdot \vec{v}_{SS'})^2 - v_{SS'}^2 c^2 (t_{SP}^2 - t_{S'P}^2) = (1 + v_{SS'}^2 c^2 b^2) (\vec{x}_{SP} \cdot \vec{v}_{SS'})^2 + 2v_{SS'}^2 c^2 a b \vec{x}_{SP} \cdot \vec{v}_{SS'} t_{SP} + (a^2 - 1) v_{SS'}^2 c^2 t_{SP}^2$$
$$\equiv [(\vec{x}_{SP} \cdot \vec{v}_{SS'}) \sqrt{1 + v_{SS'}^2 c^2 b^2} - \sqrt{a^2 - 1} c v_{SS'} t_{SP}]^2$$
(10)

which gives  $a^2 = 1 + v_{SS'}^2 c^2 b^2$ <sup>6</sup>. We have then

$$\tau = \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'} \pm \left| \vec{x}_{SP} \cdot \vec{v}_{SS'} |a| - \sqrt{a^2 - 1} c v_{SS'} t_{SP} \right|}{v_{SS'}^2} \,. \tag{11}$$

Here, the choice for the sign of  $\tau$  is fixed as follows. Replacing  $\tau$  given by (11) into (7), and considering the time transformation (9) we assume the transformation  $(t_{SP}, \vec{x}_{SP}) \rightarrow (t_{S'P}, \vec{x}_{S'P})$ 

<sup>&</sup>lt;sup>6</sup>We could have written equation (10) in the form  $(\vec{x}_{SP} \cdot \vec{v}_{SS'})\sqrt{1 + v_{SS'}^2 c^2 b^2} + \sqrt{a^2 - 1} c v_{SS'} t_{SP}$ . As we will see, the minus sign adopted in (10) becomes necessary if we intend to obtain the standard Lorentz transformation of SR.

is invertible upon replacing  $(t_{S'P}, \vec{x}_{S'P}, \vec{v}_{S'S}) \leftrightarrow (t_{SP}, \vec{x}_{SP}, \vec{v}_{SS'})$  with  $\vec{v}_{S'S} = -\vec{v}_{SS'}$ . Therefore, we notice that:

• If  $\vec{x}_{SP} \cdot \vec{v}_{SS'} |a| - \sqrt{a^2 - 1} c v_{SS'} t_{SP} \ge 0$  we must choose

$$\tau = \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'} - \left| \vec{x}_{SP} \cdot \vec{v}_{SS'} |a| - \sqrt{a^2 - 1} c v_{SS'} t_{SP} \right|}{v_{SS'}^2} = (1 - |a|) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} + \frac{\sqrt{a^2 - 1}}{v_{SS'}} c t_{SP} .$$

• If  $\vec{x}_{SP} \cdot \vec{v}_{SS'} |a| - \sqrt{a^2 - 1} c v_{SS'} t_{SP} < 0$  we must choose

$$\tau = \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'} + \left| \vec{x}_{SP} \cdot \vec{v}_{SS'} |a| - \sqrt{a^2 - 1} c v_{SS'} t_{SP} \right|}{v_{SS'}^2} = (1 - |a|) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} + \frac{\sqrt{a^2 - 1}}{v_{SS'}} c t_{SP} .$$

In both cases, we obtain the transformation

$$\begin{cases} \vec{x}_{S'P} = \vec{x}_{SP} - (1 - |a|) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} \vec{v}_{SS'} - \frac{\sqrt{a^2 - 1}}{v_{SS'}} c t_{SP} \vec{v}_{SS'} \\ t_{S'P} = |a| t_{SP} - \frac{\sqrt{a^2 - 1}}{v_{SS'}} \vec{x}_{SP} \cdot \vec{v}_{SS'} \end{cases}$$
(12)

$$\tau := (1 - |a|) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} + \frac{\sqrt{a^2 - 1}}{v_{SS'}} c t_{SP} .$$
(13)

As we have expected, the transformation (12) satisfy the condition  $\vec{x}_{S'P}^2 - c^2 t_{S'P}^2 = \vec{x}_{SP}^2 - c^2 t_{SP}^2$ , and leave  $\tau$  invariant, i.e.

$$\tau' = (1 - |a|) \frac{\vec{x}_{S'P} \cdot \vec{v}_{S'S}}{v_{S'S}^2} + \frac{\sqrt{a^2 - 1}}{v_{S'S}} c t_{S'P} = (1 - |a|) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} + \frac{\sqrt{a^2 - 1}}{v_{SS'}} c t_{SP} = \tau$$
(14)

since it was built in order to satisfy those assumptions. Here, we assume the parameter a depends on the relative speed  $v_{SS'}$ , which for convenience we denote by  $a_{v_{SS'}}$ . Under this assumption, we also obtain that  $a_{v_{S'S}} = a_{v_{SS'}}$ .

The transformation (12) depends on the velocity  $\vec{v}_{SS'}$  that is defined as the derivative relative to the absolute time

$$\vec{v}_{SS'} = \frac{d\vec{x}_{SS'}}{d\tau} \,. \tag{15}$$

Now, considering the physical time  $t_{SS'}$  we can define another velocity

$$\vec{\tilde{v}}_{SS'} := \frac{d\vec{x}_{SS'}}{dt_{SS'}} \,. \tag{16}$$

We obtain a relation between  $\vec{v}_{SS'}$ ,  $\vec{v}_{SS'}$  as follows. Identifying P with the origin of the frame S' we have from (13)

$$\tau = \frac{\sqrt{a_{v_{SS'}}^2 - 1}}{|a_{v_{SS'}}|} \frac{c}{v_{SS'}} t_{SS'}$$

which allows us to write

$$\vec{x}_{SS'} = \vec{v}_{SS'} \frac{\sqrt{a_{v_{SS'}}^2 - 1}}{|a_{v_{SS'}}|} \frac{c}{v_{SS'}} t_{SS'}$$

and then

$$\vec{\tilde{v}}_{SS'} = \vec{v}_{SS'} \frac{\sqrt{a_{v_{SS'}}^2 - 1}}{|a_{v_{SS'}}|} \frac{c}{v_{SS'}} .$$
(17)

Since the parameter  $a_{v_{SS'}}$  depends on  $v_{SS'}$  we obtain that  $\tilde{v}_{SS'}$  also depends on  $v_{SS'}$  through the expression

$$\frac{\widetilde{v}_{SS'}^2}{c^2} = \frac{a_{v_{SS'}}^2 - 1}{a_{v_{SS'}}^2} \tag{18}$$

which gives

$$\widetilde{v}_{SS'} < c . \tag{19}$$

Therefore, in our formalism every speed calculated as the derivative relative to the physical time is always less than c. Since in SR this is the type of derivative we consider, it seems the problem of tachyons is ruled out in SR. From (18) we may express  $a_{v_{SS'}}$  in terms of  $\tilde{v}_{SS'}$  as

$$a_{v_{SS'}} = \frac{1}{\sqrt{1 - \frac{\tilde{v}_{SS'}^2}{c^2}}} \equiv \gamma_{\tilde{v}_{SS'}} \tag{20}$$

and the transformation (12) becomes

$$\begin{cases} \vec{x}_{S'P} = \vec{x}_{SP} - (1 - \gamma_{\tilde{v}_{SS'}}) \frac{\vec{x}_{SP'} \cdot \vec{v}_{SS'}}{\tilde{v}_{SS'}^2} \vec{v}_{SS'} - \gamma_{\tilde{v}_{SS'}} t_{SP} \vec{\tilde{v}}_{SS'} \\ t_{S'P} = \gamma_{\tilde{v}_{SS'}} \left( t_{SP} - \frac{\vec{x}_{SP'} \cdot \vec{v}_{SS'}}{c^2} \right) \end{cases}$$
(21)

that is the familiar Lorentz transformation. Hence, we have obtained a complete equivalence between the Galilei and the Lorentz transformation.

From (17) we obtain a similar expression relating  $\vec{\tilde{v}}_{S'S}$  and  $\vec{v}_{S'S}$ , i.e.  $\frac{\tilde{v}_{S'S}^2}{c^2} = \frac{a_{v_{S'S}}^2 - 1}{a_{v_{S'S}}^2}$ , and since  $a_{v_{S'S}}$  depends on the absolute value  $|\vec{v}_{S'S}| = |\vec{v}_{SS'}|$  we end up with

$$\widetilde{v}_{S\,'S} = \widetilde{v}_{SS\,'} \; .$$

From (17) we could also have considered  $\tilde{\vec{v}}_{SS'}$  as a vector in  $\mathbb{R}^3_{SS'}$ . However, we avoid this interpretation as it would lead to an inconsistency associated to the so-called Thomas precession that we will analyze in section 3.4.3.

#### **3.3** On the role of the parameter *a*

As we have seen, the parameter a was introduced in equation (9) as a free parameter in the sense it was assumed to depend only on the relative speed  $v_{SS'}$  between the frames. The assumptions give no prescription on how to fix this dependence. Having established the relation between the absolute and the physical time, we introduced another velocity  $\tilde{v}_{SS'}$ , considered as a derivative relative to the physical time, which is related to  $\tilde{v}_{SS'}$  and the parameter  $a_{v_{SS'}}$  through equations (17, 18). Whatever the form we may take for  $a_{v_{SS'}}$ , once we obtain  $\tilde{v}_{SS'}$  as a function of  $v_{SS'}$  we may rewrite  $a_{v_{SS'}}$  in terms of  $\tilde{v}_{SS'}$ , which results in  $a_{v_{SS'}} = \gamma_{\tilde{v}_{SS'}}$  (20). This gives a "universal" character for the Lorentz transformation with the transformation (12) becoming (21). Despite this, the constraint between  $\vec{v}_{SS'}$  and  $\vec{\tilde{v}}_{SS'}$  arising from a particular choice of  $a_{v_{SS'}}$  may signalize different behaviors.

In fact, let us assume  $v_{SS'} > c$  and choose a such that

$$a_{v_{SS'}} \equiv \widetilde{\gamma} := rac{1}{\sqrt{1 - rac{c^2}{v_{SS'}^2}}}$$
 .

With this choice, the transformation (12) becomes

$$\begin{cases} \vec{x}_{S'P} = \vec{x}_{SP} - (1 - \tilde{\gamma}) \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2} \vec{v}_{SS'} - \tilde{\gamma} \frac{c^2}{v_{SS'}^2} t_{SP} \vec{v}_{SS'} \\ t_{S'P} = \tilde{\gamma} (t_{SP} - \frac{\vec{x}_{SP} \cdot \vec{v}_{SS'}}{v_{SS'}^2}) \end{cases}$$
(22)

that agrees with the expression obtained originally by Shankara (see eqs. (16),(17) of [9]) and also by Duffey [10] in the description of superluminal frames. From (18) we obtain

$$\widetilde{v}_{SS'} = \frac{c^2}{v_{SS'}} , \qquad (23)$$

a relation that is also present in the tachyonic model of [9], whose context is of a wave propagating in a medium with the velocities being interpreted in a different way, namely  $\vec{\tilde{v}}_{SS'}$  is the group velocity of the wave, while  $\vec{v}_{SS'}$  is the phase velocity. Here, even though the transformation (22) reduces to the form given on (21), which has the usual Lorentz form, the physical distinction between  $\vec{v}_{SS'}$  and  $\vec{\tilde{v}}_{SS'}$  may be employed to select one of the forms of the transformation as being physically relevant for the problem one is analyzing. We return to this issue in section 4.3.2.

### 3.4 The velocity transformation

Following the program of deducing transformations directly from the assumptions given in I-IV and the results previously obtained, we now intend to obtain the law of velocity transformation and analyze some of its consequences.

### 3.4.1 Obtaining the transformation from the Galilean velocity law

Let us assume two frames S and S' moving with velocity  $\vec{v}_{SS'}$  and such that at  $\tau = 0$  (or equivalently at t = 0 = t') both origins coincide. The position of a moving particle is written as  $\vec{x}_{SP}(\tau)$ ,  $\vec{x}_{S'P}(\tau)$ , and from  $\vec{x}_{S'P} = \vec{x}_{SP} - \vec{v}_{SS'}\tau$  we obtain the rule of velocity addition in Galilei relativity

$$\frac{d\vec{x}_{S'P}}{d\tau} = \frac{d\vec{x}_{SP}}{d\tau} - \vec{v}_{SS'} .$$

$$\tag{24}$$

From (13), and using the chain rule, we have

$$\frac{\frac{d\vec{x}_{S'P}}{dt_{S'P}}}{-(1-|a_{v_{SS'}}|)\frac{d\vec{x}_{S'P}}{dt_{S'P}} \cdot \frac{\vec{v}_{SS'}}{v_{SS'}^2} + \frac{\sqrt{a_{v_{SS'}}^2 - 1}}{v_{SS'}}c}{(1-|a_{v_{SS'}}|)\frac{d\vec{x}_{SP}}{dt_{SP}} \cdot \frac{\vec{v}_{SS'}}{v_{SS'}^2} + \frac{\sqrt{a_{v_{SS'}}^2 - 1}}{v_{SS'}}c}{(25)}$$

Let us denote

$$\vec{\tilde{v}}_{SP} \equiv \frac{d\vec{x}_{SP}}{dt_{SP}}, \quad \vec{\tilde{v}}_{S'P} \equiv \frac{d\vec{x}_{S'P}}{dt_{S'P}}$$

Taking the scalar product by  $\vec{v}_{SS'}$  on both sides of (25) we obtain

$$\vec{\tilde{v}}_{S'P} \cdot \vec{v}_{SS'} = c \, v_{SS'} \frac{|a_{v_{SS'}}| \vec{\tilde{v}}_{SP} \cdot \vec{v}_{SS'} - \sqrt{a_{v_{SS'}}^2 - 1} c \, v_{SS'}}{-\sqrt{a_{v_{SS'}}^2 - 1} \vec{\tilde{v}}_{SP} \cdot \vec{v}_{SS'} + |a_{v_{SS'}}| c \, v_{SS'}}}$$

Using this last expression back in (25) together with (17) and (18) we obtain the final form

$$\vec{\tilde{v}}_{S'P} = \frac{\vec{\tilde{v}}_{SP} - \gamma_{\tilde{v}_{SS'}} \vec{\tilde{v}}_{SS'} - \left(1 - \gamma_{\tilde{v}_{SS'}}\right) \frac{\vec{\tilde{v}}_{SP'} \vec{\tilde{v}}_{SS'}}{\vec{\tilde{v}}_{SS'}^2} \vec{\tilde{v}}_{SS'}}{\gamma_{\tilde{v}_{SS'}} \left(1 - \frac{\vec{\tilde{v}}_{SP'} \vec{\tilde{v}}_{SS'}}{c^2}\right)} .$$
(26)

Since  $\vec{v}_{S'P}$  and  $\vec{v}_{SP}$  are derivatives of the position vectors  $\vec{x}_{S'P}$ ,  $\vec{x}_{SP}$  with respect to the physical times  $t_{S'P}$  and  $t_{SP}$ , we see (26) as the "physical time" counterpart of the rule of velocity addition of the galilean relativity. It is straightforward to obtain that

$$\frac{\tilde{v}_{S'P}^2}{c^2} = 1 + \frac{\frac{\tilde{v}_{SP}^2 - 1}{c^2} - 1}{\gamma_{\tilde{v}_{SS'}} \left(1 - \frac{\tilde{v}_{SP'} \cdot \tilde{v}_{SS'}}{c^2}\right)^2}$$
(27)

and we conclude that  $\tilde{v}_{SP} > c \Rightarrow \tilde{v}_{S'P} > c$ , and  $\tilde{v}_{SP} < c \Rightarrow \tilde{v}_{S'P} < c$ . However, for a particle moving with a constant velocity relative to the frames S, S' we also have similar relations as the ones given in (18),

$$\widetilde{v}_{SP} < c , \quad \widetilde{v}_{S'P} < c ,$$

$$\tag{28}$$

therefore, the only possibility for having  $\tilde{v}_{SP} > c$ ,  $\tilde{v}_{S'P} > c$  is related to a situation that does not demand relation (18), perhaps the case of an accelerated particle that would result in a relation for  $\tau$  different from the one obtained in (13) that is the base for deducing (18).

### 3.4.2 The light speed as a derivative relative to $\tau$

Let us assume the propagation of a light ray with  $\vec{x}_{SP}$ ,  $\vec{x}_{S'P}$  being the position of a point P in the wave front as seen by S and S', i.e

$$\frac{d\vec{x}_{SP}}{dt_{SP}} = \vec{c}, \quad \frac{d\vec{x}_{S'P}}{dt_{S'P}} = \vec{c}'$$

with  $|\vec{c}'| = |\vec{c}| = c$ . We then have

$$\frac{d\vec{x}_{SP}}{d\tau} = \frac{d\vec{x}_{SP}}{dt_{SP}} \frac{dt_{SP}}{d\tau} \implies c_S = \frac{c}{\left|(1 - |a|)\frac{\vec{c}\cdot\vec{v}_{SS'}}{v_{SS'}^2} + \sqrt{a^2 - 1}\frac{c}{v_{SS'}}\right|}$$
(29)  
$$\frac{d\vec{x}_{S'P}}{d\tau} = \frac{d\vec{x}_{S'P}}{dt_{S'P}} \frac{dt_{S'P}}{d\tau} \implies c_{S'} = \frac{c}{\left|-(1 - |a|)\frac{\vec{c}'\cdot\vec{v}_{SS'}}{v_{SS'}^2} + \sqrt{a^2 - 1}\frac{c}{v_{SS'}}\right|}$$

that implies different values for the speed when considered as a rate of change relative to the absolute time, i.e.

$$c_S = \left| \frac{d\vec{x}_{SP}}{d\tau} \right| \neq c'_S = \left| \frac{d\vec{x}_{S'P}}{d\tau} \right|$$

Here, there is one more consistency check to be performed. We expect to have the following relation  $|\vec{x}_{SP}| = c_S \Delta \tau = c \Delta t$ . In fact, consider an event  $P : (t, \vec{x}) \sim (\tau, \vec{x})$  as described by S. At this instant, assume the observer in the frame S that is at the same location of the event Psends a light ray towards the observer that is at the origin of S. The arrival of the light ray at the origin of S corresponds to another event described by S as  $(t_1, 0) \sim (\tau_1, 0)$ . Here, in order to relate t and  $\tau$ , and  $t_1$  and  $\tau_1$  we assume any auxiliary frame S' moving relative to S with velocity  $\vec{v}$  and employ (14). For the events  $(t, \vec{x}) \sim (\tau, \vec{x}), (t_1, 0) \sim (\tau_1, 0)$  we then have

$$\tau = (1 - |a|)\frac{\vec{x} \cdot \vec{v}}{v^2} + \frac{\sqrt{a^2 - 1}}{v} ct$$
$$\tau_1 = \frac{\sqrt{a^2 - 1}}{v} ct_1 .$$

Denoting  $\Delta t := t_1 - t$  and expressing  $\vec{x} = -\vec{c}\Delta t$  we obtain

$$\Delta \tau := \tau_1 - \tau = \left[\sqrt{a^2 - 1}\frac{c}{v} + (1 - |a|)\frac{\vec{c} \cdot \vec{v}}{v^2}\right] \Delta t \; .$$

Then,

$$|\vec{x}| = c\,\Delta t = \frac{c}{\sqrt{a^2 - 1}\frac{c}{v} + (1 - |a|)\frac{\vec{c}\cdot\vec{v}}{v^2}}\,\,\Delta\tau$$

and from (29) we obtain  $|\vec{x}| = c_S \Delta \tau$ .

### 3.4.3 The composition of velocities

Let us assume three frames  $\Sigma, S, S'$  moving relative to each other with velocities  $\vec{v}_{\Sigma S}, \vec{v}_{\Sigma S'}, \vec{v}_{S S'}$ , having their axes parallel and their origins coinciding at  $\tau = 0$ . According to Galilei relativity we write

$$\vec{x}_{\Sigma S'} = \vec{x}_{\Sigma S} + \vec{x}_{S S'} \tag{30}$$

with

$$\vec{x}_{\Sigma S'} = \vec{v}_{\Sigma S'} \tau , \quad \vec{x}_{\Sigma S} = \vec{v}_{\Sigma S} \tau , \quad \vec{x}_{S S'} = \vec{v}_{S S'} \tau .$$

$$(31)$$

In particular, we have  $\vec{x}_{\Sigma S'} = -\vec{x}_{S'\Sigma}$ , which results

$$\vec{v}_{\Sigma S'} = -\vec{v}_{S'\Sigma} , \qquad (32)$$

We now analyze the movement of S' relative to S and  $\Sigma$ , and then the movement of  $\Sigma$  relative to S and S' using the same procedure that led to (26).

# (i) The movement of S' relative to $\Sigma$ and S

For the movement of S' relative to  $\Sigma$  and S we have  $(t_{\Sigma S'}, \vec{x}_{\Sigma S'}) \sim (t_{SS'}, \vec{x}_{SS'})$  and

$$\tau = \begin{cases} (1 - |a_{v_{\Sigma S}}|) \frac{\vec{x}_{SS'} \cdot (-\vec{v}_{\Sigma S})}{v_{\Sigma S}^2} + \frac{\sqrt{a_{v_{\Sigma S}}^2 - 1}}{v_{\Sigma S}} c \, t_{SS'} \\ (1 - |a_{v_{\Sigma S}}|) \frac{\vec{x}_{\Sigma S'} \cdot \vec{v}_{\Sigma S}}{v_{\Sigma S}^2} + \frac{\sqrt{a_{v_{\Sigma S}}^2 - 1}}{v_{\Sigma S}} c \, t_{\Sigma S'} \,. \end{cases}$$
(33)

From (30) we obtain

$$\frac{d\vec{x}_{\Sigma S'}}{dt_{\Sigma S'}}\frac{dt_{\Sigma S'}}{d\tau} = \vec{v}_{\Sigma S} + \frac{d\vec{x}_{S S'}}{dt_{S S'}}\frac{dt_{S S'}}{d\tau} .$$

$$(34)$$

Denoting

$$\vec{\tilde{v}}_{\Sigma S'} \equiv \frac{d\vec{x}_{\Sigma S'}}{dt_{\Sigma S'}}, \ \vec{\tilde{v}}_{S S'} \equiv \frac{d\vec{x}_{S S'}}{dt_{S S'}}$$

and using (33), we obtain by the same procedure employed in obtaining (26) the following expression

$$\vec{\tilde{v}}_{\Sigma S'} = \frac{\vec{\tilde{v}}_{S S'} + \gamma_{\tilde{v}_{\Sigma S}} \vec{\tilde{v}}_{\Sigma S} - \left(1 - \gamma_{\tilde{v}_{\Sigma S}}\right) \frac{\vec{\tilde{v}}_{S S'} \cdot \vec{\tilde{v}}_{\Sigma S}}{\vec{\tilde{v}}_{\Sigma S}^2} \vec{\tilde{v}}_{\Sigma S}}{\gamma_{\tilde{v}_{\Sigma S}} \left(1 + \frac{\vec{\tilde{v}}_{S S'} \cdot \vec{\tilde{v}}_{\Sigma S}}{c^2}\right)}$$
(35)

## (ii) The movement of $\Sigma$ relative to S and S'

For the movement of  $\Sigma$  relative to S and S' we have  $(t_{S\Sigma}, \vec{x}_{S\Sigma}) \sim (t_{S'\Sigma}, \vec{x}_{S'\Sigma})$  and

$$\tau = \begin{cases} (1 - |a_{v_{S'S}}|) \frac{\vec{x}_{S\Sigma} \cdot (-\vec{v}_{S'S})}{v_{S'S}^2} + \frac{\sqrt{a_{v_{S'S}}^2 - 1}}{v_{S'S}} c t_{S\Sigma} \\ (1 - |a_{v_{S'S}}|) \frac{\vec{x}_{S'\Sigma} \cdot \vec{v}_{S'S}}{v_{S'S}^2} + \frac{\sqrt{a_{v_{S'S}}^2 - 1}}{v_{S'S}} c t_{S'\Sigma} . \end{cases}$$
(36)

From (30) and using that  $\vec{x}_{S'\Sigma} = -\vec{x}_{\Sigma S'}, \ \vec{x}_{\Sigma S} = -\vec{x}_{S\Sigma}, \ \vec{x}_{S'S} = -\vec{x}_{SS'}$  we now have

$$\frac{d\vec{x}_{S'\Sigma}}{dt_{S'\Sigma}}\frac{dt_{S'\Sigma}}{d\tau} = \frac{d\vec{x}_{S\Sigma}}{dt_{S\Sigma}}\frac{dt_{S\Sigma}}{d\tau} + \vec{v}_{S'S} .$$
(37)

Denoting

$$\vec{\tilde{v}}_{S'\Sigma} \equiv \frac{d\vec{x}_{S'\Sigma}}{dt_{S'\Sigma}} , \quad \vec{\tilde{v}}_{S\Sigma} \equiv \frac{d\vec{x}_{S\Sigma}}{dt_{S\Sigma}}$$
(38)

we obtain

$$\vec{\tilde{v}}_{S'\Sigma} = \frac{\vec{\tilde{v}}_{S\Sigma} + \gamma_{\tilde{v}_{SS'}} \vec{\tilde{v}}_{S'S} - (1 - \gamma_{\tilde{v}_{SS'}}) \frac{\tilde{v}_{S\Sigma'} \cdot \tilde{v}_{S'S}}{\tilde{v}_{S'S}^2} \vec{\tilde{v}}_{S'S}}{\gamma_{\tilde{v}_{SS'}} (1 + \frac{\vec{\tilde{v}}_{S\Sigma} \cdot \vec{\tilde{v}}_{S'S}}{c^2})},$$
(39)

where we have used that  $\tilde{v}_{SS'} = \tilde{v}_{S'S}$  to identify  $\gamma_{\tilde{v}_{SS'}} = \gamma_{\tilde{v}_{S'S}}$ . We recognize equations (35, 39) with the velocity transformations obtained by C. Moller in his analysis of the Thomas precession [5]. As we know, in this particular case, the velocities  $\vec{v}_{\Sigma S'}$ ,  $\vec{v}_{S'\Sigma}$  are related by a rotation [5, 11]. We notice that each of the equations (35, 39) involve vectors belonging to different spaces, therefore we should interpret each equation in the same way as we did for equation (5), i.e.  $\vec{v}_{\Sigma S'}, \vec{v}_{S'\Sigma}$  are vectors belonging to an abstract space V where the equations (35, 39) make sense. It is in this space that we have

$$\vec{\tilde{v}}_{\Sigma S'} \neq -\vec{\tilde{v}}_{S'\Sigma} , \qquad (40)$$

which accounts for the Thomas precession. However, the vectors  $\vec{v}_{\Sigma S'}, \vec{v}_{S'\Sigma}$  are defined in the space  $\mathbb{R}^3_{\Sigma SS'}$  where they satisfy (32), which is a consequence of the frames  $\Sigma, S, S'$  having their axis parallel during its translational movement. Now, it becomes clear the reason for not identifying the velocities  $\vec{v}_{\Sigma S'}, \vec{v}_{S'\Sigma}$  as vectors in  $\mathbb{R}^3_{\Sigma SS'}$ . Indeed, if this were the case, then using (17) we would associate each of them with the vectors  $\vec{v}_{\Sigma S'}, \vec{v}_{S'\Sigma}$  through

$$\vec{\tilde{v}}_{\Sigma S'} = \vec{v}_{\Sigma S'} \frac{\sqrt{a_{v_{\Sigma S'}}^2 - 1}}{|a_{v_{\Sigma S'}}|} \frac{c}{v_{\Sigma S'}} , \qquad \vec{\tilde{v}}_{S'\Sigma} = \vec{v}_{S'\Sigma} \frac{\sqrt{a_{v_{S'\Sigma}}^2 - 1}}{|a_{v_{S'\Sigma}}|} \frac{c}{v_{S'\Sigma}}$$

and from (32) we would have  $\vec{\tilde{v}}_{\Sigma S'} = -\vec{\tilde{v}}_{S'\Sigma}$ , which contradicts (40).

### 3.5 The structure of space and time

Spacetime is the physical arena where events occur. Therefore, in order to describe spacetime we must endow it with a suitable coordinate system, e.g. the one provided by a reference frame, which is how we describe events. In our approach, we have an equivalence between the absolute and the physical time that allows an observer to associate coordinates to events according to  $(\tau, \vec{x})$  or  $(t, \vec{x})$ , and from each of these coordinatizations originates a particular view for the spacetime. In what follows, we search for a description of spacetime that combines the views of the Galilei and the Special relativity.

Given two frames S, S' let us write the Galilei transformation as

$$(\tau_S, \vec{x}_S) \to (\tau_{S'}, \vec{x}_{S'}) : \begin{cases} \tau_S = \tau_{S'} \equiv \tau \\ \vec{x}_{S'} = \vec{x}_S - \vec{v}\tau \end{cases}$$

Due to the absolute character of  $\tau$  we have that for a fixed  $\tau \equiv \tau_S = \tau_{S'}$  the Galilei transformation is essentially a map  $\mathbb{R}^3_S \to \mathbb{R}^3_{S'}$ , where  $\mathbb{R}^3_S \ni \vec{x}_S, \mathbb{R}^3_{S'} \ni \vec{x}_{S'}$  denote the 3-dimensional Euclidean space. When spacetime is described using  $(\tau, \vec{x})$  as coordinates we refer to it as the *Galilei* spacetime, that admits the following representation  $\cup_{\tau \in \mathbb{R}} (\{\tau\} \times \mathbb{R}^3) \equiv \mathbb{R} \times \mathbb{R}^3$ . In this sense, the Galilei spacetime decomposes in terms of 3-dimensional spaces  $\mathbb{R}^3$  that are indexed by  $\tau \in \mathbb{R}$ .

Now, let us analyze the transformation under the form (12). For the same fixed value of  $\tau$ , we rewrite equation (14) as

$$\tau = x^0 n^0 + \vec{x} \cdot \vec{n} = x'^0 n'^0 + \vec{x}' \cdot \vec{n}' \tag{41}$$

with

$$n^{\mu} \equiv (n^{0}, \vec{n}) := \left(\frac{\sqrt{a^{2} - 1}}{v}, -\frac{(|a| - 1)\vec{v}}{v^{2}}\right)$$

$$n^{\prime \mu} \equiv (n^{\prime 0}, \vec{n}^{\prime}) := \left(\frac{\sqrt{a^{2} - 1}}{v}, \frac{(|a| - 1)\vec{v}}{v^{2}}\right)$$
(42)

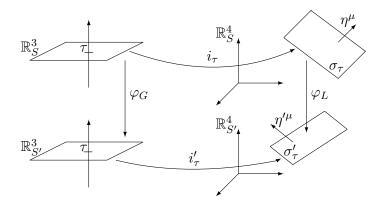
and  $x^{\mu} \equiv (x^0, \vec{x}) := (ct, \vec{x}), \ x'^{\mu} \equiv (x'^0, \vec{x}') := (ct', \vec{x}')$ . We identify (41) as the equation of two hyperplanes  $\sigma_{\tau} : \tau = x^0 n^0 + \vec{x} \cdot \vec{n}$  and  $\sigma'_{\tau} : \tau = x'^0 n'^0 + \vec{x}' \cdot \vec{n}'$  defined respectively in the spaces  $\mathbb{R}^4_S$  and  $\mathbb{R}^4_{S'}$  (the spaces  $\mathbb{R}^4_S \ni (ct, \vec{x}), \mathbb{R}^4_{S'} \ni (ct', \vec{x}')$  serve here as the background in terms of which each observer describes spacetime). Then, for a fixed value of  $\tau$  the transformation (12) is seen as a transformation  $\sigma_{\tau} \to \sigma'_{\tau}$  between these hyperplanes. Here, we refer the spacetime as the *SR spacetime* and it decomposes as  $\cup_{\tau \in \mathbb{R}} \sigma_{\tau} \equiv \mathbb{R}^4$ .

From the previous considerations the properties of the spacetime that arise from our unified model are summarized as follows. First, we notice that the Galilei and SR spacetime refers to the same spacetime, in fact, they are just different descriptions of the same structure. Second, both the Galilei and the SR spacetime are sliced by 3-dimensional spaces, the Galilei spacetime being "sliced" by hyperplanes that are normal to the  $\tau$ -direction or, equivalently, by planes that correspond to the space  $\mathbb{R}^3$ . For the SR spacetime the "slicing" is made by hyperplanes  $\sigma_{\tau}$ ,  $\sigma'_{\tau}$ that are orthogonal to the direction of the vectors n and n'. In particular, the relation between the Galilei transformation and the Lorentz transformation is clarified by the diagram below,

where for a fixed  $\tau$  and relative to two frames of reference S, S' that move with relative velocity  $\vec{v}$  we have defined mappings

$$\begin{cases} i_{\tau} : \mathbb{R}^3_S \to \mathbb{R}^4_S \\ \vec{x} \to (ct(\tau, \vec{x}), \vec{x}) \\ t(\tau, \vec{x}) = \frac{v}{c\sqrt{a^2 - 1}} \Big( \tau + (|a| - 1)\frac{\vec{x} \cdot \vec{v}}{v^2} \Big) \end{cases} \qquad \begin{cases} i'_{\tau} : \mathbb{R}^3_{S'} \to \mathbb{R}^4_{S'} \\ \vec{x}' \to (ct'(\tau, \vec{x}'), \vec{x}') \\ t'(\tau, \vec{x}') = \frac{v}{c\sqrt{a^2 - 1}} \Big( \tau - (|a| - 1)\frac{\vec{x}' \cdot \vec{v}}{v^2} \Big) . \end{cases}$$

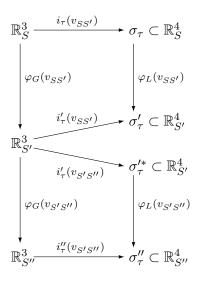
We notice that endowing  $\mathbb{R}^3_S$ ,  $\mathbb{R}^3_{S'}$ ,  $\mathbb{R}^4_S$ ,  $\mathbb{R}^4_{S'}$  with the standard Euclidean topology we obtain  $i_{\tau}(\mathbb{R}^3_S) = \sigma_{\tau}$ ,  $i'_{\tau}(\mathbb{R}^3_{S'}) = \sigma'_{\tau}$  as embedded respectively in  $\mathbb{R}^4_S$  and  $\mathbb{R}^4_{S'}$ . Then, given a Galilei transformation  $(\tau, \vec{x}) \xrightarrow{\varphi_G} (\tau, \vec{x}')$  the embeddings  $i_{\tau}, i'_{\tau}$  transfer  $\varphi_G$  to the spaces  $\mathbb{R}^4_S$ ,  $\mathbb{R}^4_{S'}$  where it assumes the form of the Lorentz transformation  $(ct, \vec{x}) \xrightarrow{\varphi_L} (ct', \vec{x}')$ , which is shown in the schema below



This view of the transformation provides an alternative way to analyze some aspects of the composition of Lorentz transformations. Indeed, let us write the Lorentz transformation as

$$\varphi_L(v_{SS'}) = i'_{\tau}(v_{SS'}) \circ \varphi_G(v_{SS'}) \circ i^{-1}_{\tau}(v_{SS'})$$
(43)

and consider three frames S, S', S'' that move with relative velocity  $\vec{v}_{SS'}, \vec{v}_{SS''}, \vec{v}_{S'S''}$ . Consider then the following diagram



We notice that since the hyperplanes  $\sigma'_{\tau}$  and  $\sigma'^*_{\tau}$  are different the Lorentz transformations  $\varphi_L(v_{SS'})$  and  $\varphi_L(v_{S'S''})$  may be composed only through the maps  $i'_{\tau}(v_{S'S''}) \circ i'^{-1}_{\tau}(v_{SS'})$ , e.g.

$$\varphi_L(v_{S'S''}) \circ i'_{\tau}(v_{S'S''}) \circ i'^{-1}_{\tau}(v_{SS'}) \circ \varphi_L(v_{SS'})$$

and we have

$$\begin{aligned} \varphi_L(v_{S'S''}) \circ i'_{\tau}(v_{S'S''}) \circ i'_{\tau}^{-1}(v_{SS'}) \circ \varphi_L(v_{SS'}) &= i''_{\tau}(v_{S'S''}) \circ \varphi_G(v_{SS'}) \circ i_{\tau}^{-1}(v_{SS'}) \\ &= i''_{\tau}(v_{S'S''}) \circ \varphi_G(v_{SS''}) \circ i_{\tau}^{-1}(v_{SS'}) \\ &\neq \varphi_L(v_{SS''}) \,. \end{aligned}$$

Therefore, from this interpretation of the Lorentz transformation acting on hyperplanes associated to the same value of  $\tau$ , we reobtain the common fact that the composition of two Lorentz transformations is not a Lorentz transformation. Here, the reason is that even though the composition  $\varphi_G(v_{S'S''}) \circ \varphi_G(v_{SS'})$  may be identified with  $\varphi_G(v_{SS''})$  the later induces the Lorentz transformation  $\varphi_L(v_{SS'})$  only through the embeddings  $i'_{\tau}(v_{SS'})$ ,  $i^{-1}_{\tau}(v_{SS'})$ , as shown in (43).

# 4 Model II

### 4.1 Assumptions

We propose another model based on the following assumptions:

### I. Space

Each inertial frame describes space as an euclidean 3-dimensional vector space.

## II. Time

Time is described by a variable t that depends on the frame, in such way that two inertial frames S and S' measure time t and t'.

**III.** The speed of light is constant relative to inertial frames.

**IV.** Relative to two inertial frames any event  $(t, \vec{x}) \sim (t', \vec{x}')$  satisfies

$$\alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta} \cdot \vec{x}' \tag{44}$$

with  $\alpha, \vec{\beta}$  arbitrary parameters.

**V.** The relation between  $(t, \vec{x})$  and  $(t', \vec{x}')$  is linear.

#### **Remarks:**

(i) Postulates I and II are essentially the same postulates I and II of model I, except that in the current postulate II of model II there is no mention to the absolute time.

(ii) The condition  $\alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta} \cdot \vec{x}'$  is suggested by a previous condition given in (41), but now with the constants  $\alpha$ ,  $\vec{\beta}$  being arbitrary parameters. Here, equation (44) is the starting point from which we will derive the transformation. In this sense, it plays the same role as the relation  $\vec{x}'^2 - c^2 t'^2 = \vec{x}^2 - c^2 t^2$  of the standard SR formulation. However, while in SR this last equation may be expressed in terms of the invariance of the quadratic form  $Q(x^{\mu}) := \vec{x}^2 - c^2 t^2$ , which is a quantity defined uniquely in terms of the spacetime coordinates, the same can't be said of the equation (44), since it contains explicitly the parameters  $\alpha$ ,  $\vec{\beta}$  that, in general, don't refer to any known property of the space and time. Therefore, unless we are able to identify the parameters  $\alpha$ ,  $\vec{\beta}$  with any intrinsic property of the spacetime it will not be possible to associate any mathematical structure to the equation (44) as it was done with  $Q(x^{\mu})$ .

### 4.2 The Generalized Lorentz Transformation

From V we write

$$\begin{split} x^{\prime 0} &= A^0_{\ 0} x^0 + A^0_{\ i} x^i \\ x^{\prime i} &= A^i_{\ 0} x^0 + A^i_{\ j} x^j \end{split}$$

From IV, we obtain by replacing them into (44) that

$$A^{0}_{\ 0} = 1 + \frac{1}{\alpha} A^{i}_{\ 0} \beta^{i} \tag{45}$$

$$A^{0}_{\ i} = \frac{1}{\alpha} \left( \beta^{i} + A^{j}_{\ i} \beta^{j} \right) \tag{46}$$

which momentarily gives

$$t' = \left(1 + \frac{1}{\alpha}A^i_0\beta^i\right)t + \frac{1}{\alpha c}\left(\beta^i + A^j_i\beta^j\right)x^i$$
  
$$x'^i = A^i_0ct + A^i_jx^j .$$
(47)

Here we assume

$$A^{i}_{\ 0} = \lambda \beta^{i}$$

$$A^{i}_{\ j} = \delta^{i}_{\ j} + \xi \beta^{i} \beta^{j}$$
(48)

with  $\lambda, \xi$  arbitrary quantities that will be fixed later. We require the inverse transformation is obtained by the change  $(t, \vec{x}, \beta^i) \leftrightarrow (t', \vec{x}', -\beta^i)$  into (47). This fixes  $\lambda = \xi \alpha$ . Let us consider now the expression for the velocity. Here, we have

$$\frac{dx'^{i}}{dt'} = \frac{A^{i}_{\ 0}c + A^{i}_{\ j}\frac{dx^{j}}{dt}}{1 + \frac{1}{\alpha}A^{i}_{\ 0}\beta^{i} + \frac{1}{\alpha c}(\beta^{i} + A^{j}_{\ i}\beta^{j})\frac{dx^{i}}{dt}}.$$
(49)

From III the speed of light is constant relative to any inertial frame. Then for a light wave we have

$$\frac{d\vec{x}'}{dt'} = \vec{c}', \quad \frac{d\vec{x}}{dt} = \vec{c} \equiv c\hat{\eta} \quad \text{with} \quad \hat{\eta} \cdot \hat{\eta} = 1 \text{ and } |\vec{c}'| = |\vec{c}|.$$

From (49), (48), using  $\lambda = \xi \alpha$ , and due to the arbitrariness of the direction of the unit vector  $\hat{\eta}$  we obtain

$$\xi = \frac{2}{\alpha^2 \left(1 - \frac{\beta^2}{\alpha^2}\right)} \tag{50}$$

which then fixes all the transformation constants  $A_0^0, A_i^0, A_0^i, A_j^i$  in terms of  $\alpha, \beta$  as follows

$$A_{0}^{0} = \frac{1 + \frac{\beta^{2}}{\alpha^{2}}}{1 - \frac{\beta^{2}}{\alpha^{2}}}, \quad A_{i}^{0} = A_{0}^{i} = \frac{2}{1 - \frac{\beta^{2}}{\alpha^{2}}} \frac{\beta^{i}}{\alpha}, \quad A_{j}^{i} = \delta_{j}^{i} + \frac{2}{1 - \frac{\beta^{2}}{\alpha^{2}}} \frac{\beta^{i}}{\alpha} \frac{\beta^{j}}{\alpha}$$
(51)

and finally leave us with the transformations

$$t' = \left(\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}}\right) \left\{ t + \frac{2}{c\left(1+\frac{\beta^2}{\alpha^2}\right)} \frac{\vec{\beta}}{\alpha} \cdot \vec{x} \right\}$$
  
$$\vec{x}' = \vec{x} + \frac{2}{1-\frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} \cdot \vec{x} \frac{\vec{\beta}}{\alpha} + \frac{2ct}{1-\frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} .$$
 (52)

We have then obtained a class of transformations that is compatible with the invariance of the speed of light and that depends on arbitrary parameters  $\alpha, \vec{\beta}$ . We call this transformation Generalized Lorentz Transformation (GLT). It is immediate to check that it satisfies  $\vec{x}'^2 - c^2 t'^2 = \vec{x}^2 - c^2 t^2$  (and not only for a light wave), which shows that it must include the standard Lorentz transformation as a particular case.

**Remark**: It is possible to work with a more general relation than the one given in (44), for example, if we assume that  $\vec{\beta}$  can be transformed then we would write

$$\alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta}' \cdot \vec{x}'$$

and the effect of the invariance of  $\vec{\beta}' \cdot \vec{x}'$  in this new condition corresponds to replace  $\vec{\beta} \to \mathcal{R}\vec{\beta}$ ,  $\vec{x} \to \mathcal{R}\vec{x}$  in (52), i.e.

$$\begin{cases} t' = \left(\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}}\right) \left\{ t + \frac{2}{c\left(1+\frac{\beta^2}{\alpha^2}\right)} \frac{\vec{\beta}}{\alpha} \cdot \vec{x} \right\} \\ \vec{x}' = \mathcal{R}\vec{x} + \frac{2}{1-\frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} \cdot \vec{x} \frac{\mathcal{R}\vec{\beta}}{\alpha} + \frac{2ct}{1-\frac{\beta^2}{\alpha^2}} \frac{\mathcal{R}\vec{\beta}}{\alpha} \end{cases}$$
(53)

with  $\mathcal{R}$  an arbitrary space rotation.

In [12] E. Kapuscik using only the assumption of the linearity of the transformation and the constancy of the speed of light proposed a generalization of the Lorentz transformation having the form

$$\begin{cases} t' = At + \vec{B} \cdot \vec{x} \\ \vec{x}' = \sqrt{A^2 - c^2 \vec{B}^2} \left(\mathcal{R}\vec{x}\right) + \frac{A - \sqrt{A^2 - c^2 \vec{B}^2}}{\vec{B}^2} \left(\mathcal{R}\vec{B}\right) \left(\vec{B} \cdot \vec{x}\right) + c^2 \left(\mathcal{R}\vec{B}\right) t . \end{cases}$$
(54)

Comparison of the time transformation given in (53) and (54) suggests the identification

$$A = \frac{1 + \frac{\beta^2}{\alpha^2}}{1 - \frac{\beta^2}{\alpha^2}}, \qquad \vec{B} = \frac{2}{c\left(1 - \frac{\beta^2}{\alpha^2}\right)}\frac{\vec{\beta}}{\alpha}$$
(55)

that also shows the equivalence of the space transformation. In our formalism, up to the derivation of (53), there is no restriction on the parameters  $\alpha, \vec{\beta}$ , while in the development of [12] they have to obey  $A^2 - c^2 \vec{B}^2 > 0$ . Since in [12] the relative velocity  $\vec{V}$  between the frames is identified with  $-c^2 \frac{\vec{B}}{A}$ , the previous inequality leads to the restriction V < c. As we will show in sections 4.3.1 and 4.3.2, in our formalism there is no restriction on the relative speed of the frames. Having obtained the GLT given in (52) we analyze in the sequence some of its properties.

### 4.3 Transformation of Velocity

Associated to GLR we obtain by replacing (51) into (49) the associated transformation of velocity as

$$\frac{d\vec{x}\,'}{dt'} = \frac{\frac{d\vec{x}}{dt} + \frac{2c}{1 - \frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} + \frac{2}{1 - \frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} \cdot \frac{d\vec{x}}{dt} \frac{\vec{\beta}}{\alpha}}{\frac{1 + \frac{\beta^2}{\alpha^2}}{1 - \frac{\beta^2}{\alpha^2}} + \frac{2}{c(1 - \frac{\beta^2}{\alpha^2})} \frac{\vec{\beta}}{\alpha} \cdot \frac{d\vec{x}}{dt}}$$
(56)

and

$$\left(\frac{d\vec{x}\,'}{dt'}\right)^2 - c^2 = \frac{\left(\frac{d\vec{x}}{dt}\right)^2 - c^2}{\left(\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}} + \frac{2}{c(1-\frac{\beta^2}{\alpha^2})}\frac{\vec{\beta}}{\alpha} \cdot \frac{d\vec{x}}{dt}\right)^2} \tag{57}$$

therefore we obtain for the GLT that

$$\left|\frac{d\vec{x}}{dt}\right| < c \Leftrightarrow \left|\frac{d\vec{x}\,'}{dt'}\right| < c \qquad \text{or} \qquad \left|\frac{d\vec{x}}{dt}\right| > c \Leftrightarrow \left|\frac{d\vec{x}\,'}{dt'}\right| > c$$

i.e. relative to two frames described by transformation (52) a particle that is described as subluminal (superluminal) particle by one frame will also be described as subluminal (superluminal) particle by the other frame.

# 4.3.1 A transformation involving subluminal frames: $\tilde{v} < c$

As a particular case of transformation (52) let us choose the parameters  $\alpha, \vec{\beta}$  in such way that

$$\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}} = \frac{1}{\sqrt{1-\frac{\widetilde{v}^2}{c^2}}}$$

with  $\tilde{v} < c$ . We obtain

$$\frac{\vec{\beta}}{\alpha} = -c \left(1 - \sqrt{1 - \frac{\widetilde{v}^2}{c^2}}\right) \frac{\vec{\widetilde{v}}}{\widetilde{v}^2}$$

that replacing in (52) leads to the transformation

$$t' = \gamma \left( t - \frac{\vec{x} \cdot \vec{\tilde{v}}}{c^2} \right)$$
$$\vec{x}' = \vec{x} - (1 - \gamma) \frac{\vec{x} \cdot \vec{\tilde{v}}}{\tilde{v}^2} \vec{\tilde{v}} - \gamma t \vec{\tilde{v}}$$

with  $\gamma:=\frac{1}{\sqrt{1-\frac{\tilde{v}^2}{c^2}}},$  that corresponds to the Lorentz transformation.

# 4.3.2 A transformation involving superluminal frames: $\tilde{v} > c$

Let us choose  $\alpha, \vec{\beta}$  satisfying

$$\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}} = \frac{1}{\sqrt{1-\frac{c^2}{\widetilde{v}^2}}}$$

with  $\tilde{v} > c$ . Now we obtain

$$\frac{\vec{\beta}}{\alpha} = -\frac{1}{c} \left( 1 - \sqrt{1 - \frac{c^2}{\tilde{v}^2}} \right) \vec{v}$$

and replacing it into (52) gives the transformation

$$t' = \widetilde{\gamma} \left( t - \frac{\vec{x} \cdot \vec{v}}{\widetilde{v}^2} \right)$$
  
$$\vec{x}' = \vec{x} - (1 - \widetilde{\gamma}) \frac{\vec{x} \cdot \vec{v}}{\widetilde{v}^2} \vec{\tilde{v}} - \widetilde{\gamma} \frac{c^2}{\widetilde{v}^2} t \vec{\tilde{v}}$$
(58)

with  $\tilde{\gamma} := \frac{1}{\sqrt{1-\frac{c^2}{\tilde{v}^2}}}$ . Here, we reobtained Shankara and Duffey's transformation for tachyons [9, 10]. This transformation (58) appeared originally in [9] in the description of the propagation of a wave in a medium with one of the frames being at rest relative to the medium and the other moving with a speed greater than c (in fact, in [9] since the medium is not the vacuum, instead of c the light speed is denoted by  $w_0$ ). The physical content of this transformation has been discussed more generally in [10], which we refer the reader for details. Here we wish to analyze the role played by the velocity  $\vec{\tilde{v}}$  in the transformation. We consider two situations.

First, let us consider the movement of an object such that according to the frame S the object is at rest relative to the frame S'. Assume  $\tilde{\vec{v}}$  is the relative velocity between the frames. Then the frame S writes for the position of the object  $\vec{x} = \tilde{\vec{v}}t$ . Replacing it in (58) we obtain t' = 0 and  $\vec{x}' = \frac{1}{\gamma}t\vec{\vec{v}} = \frac{1}{\gamma}\vec{x}$ . Therefore, for an interval  $(0, \Delta t)$ , S associates to the movement of the object the set of events  $\{(t, \vec{x}) := (t, \tilde{\vec{v}}t) : 0 \le t \le \Delta t\}$ , while S' associates the set of events  $\{(t', \vec{x}') = (0, \frac{1}{\gamma}\vec{x}) : \vec{x} = \tilde{\vec{v}}t\}$ . Since the whole movement of the object as described by S reduces in S' to a static situation where t' = 0, and since there are many  $\vec{x}'$  corresponding to the many possible values of  $\vec{x}$ , we conclude that relative to S' the object seems to have a spatial extension. Therefore, from the perspective of the frame S' the object cannot be seen as a particle in the classical context. This is reinforced by noticing that the possibility of considering  $\vec{x} = \vec{\tilde{v}t}$  leads to the vanishing of the denominator of (56) that shows the velocity transformation law does not apply.

Let us consider now the movement of an object such that according to the frame S' the object is at rest relative to the frame S', i.e.  $\vec{x}' = 0$ . Replacing it in (58) we now obtain  $t' = \frac{1}{\tilde{\gamma}}t$ , and  $\vec{x} = \frac{c^2}{\tilde{v}^2}\tilde{v}t$ . Therefore we get

$$\left|\frac{d\vec{x}}{dt}\right| = \frac{c^2}{\widetilde{v}} \;,$$

and being  $\tilde{v} > c$  we end up with  $\left|\frac{d\vec{x}}{dt}\right| < c$ . Here, it is natural that instead of  $\vec{\tilde{v}}$  we assume the relative velocity between the frames to be  $\vec{V} := \frac{c^2}{\tilde{v}^2} \vec{\tilde{v}}$ , in terms of which the transformation (58) becomes

$$t' = \gamma \left( t - \frac{\vec{x} \cdot \vec{V}}{V^2} \right)$$
$$\vec{x}' = \vec{x} - (1 - \gamma) \frac{\vec{x} \cdot \vec{V}}{V^2} \vec{V} - \gamma t \vec{V}$$

with  $\gamma = \frac{1}{\sqrt{1 - \frac{V^2}{c^2}}}$ . Then, from a theoretical perspective, we conclude that a transformation like (58) may refer to different situations according to the interpretation one gives to  $\vec{\tilde{v}}$ .

### 4.3.3 Another superluminal transformation

As another class of transformation let us choose  $\alpha, \vec{\beta}$  such that

$$\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}} = \frac{1}{\sqrt{\frac{\widetilde{v}^2}{c^2}-1}}$$

with  $\tilde{v} > c$ , which gives

$$\frac{\vec{\beta}}{\alpha} = -\sqrt{\frac{1-\sqrt{\frac{\tilde{v}^2}{c^2}-1}}{1+\sqrt{\frac{\tilde{v}^2}{c^2}-1}}} \ \frac{\vec{v}}{\tilde{v}} \ .$$

Replacing it in (52) we obtain the transformation

$$\begin{aligned} t' &= \frac{1}{\sqrt{\frac{\widetilde{v}^2}{c^2} - 1}} \bigg\{ t - \sqrt{2 - \frac{\widetilde{v}^2}{c^2}} \, \frac{\vec{x} \cdot \vec{v}}{c \widetilde{v}} \bigg\} \\ \vec{x}' &= \vec{x} + \frac{1 - \sqrt{\frac{\widetilde{v}^2}{c^2} - 1}}{\sqrt{\frac{\widetilde{v}^2}{c^2} - 1}} \, \frac{\vec{x} \cdot \vec{v}}{\widetilde{v}^2} \, \vec{v} - \frac{\sqrt{2 - \frac{\widetilde{v}^2}{c^2}}}{\sqrt{\frac{\widetilde{v}^2}{c^2} - 1}} \, \frac{c}{\widetilde{v}} t \vec{v} \end{aligned}$$

The condition of having real transformations requires that  $\tilde{v} < 2c$ , which becomes an upper bound velocity for tachyonic motion.

# 4.4 Introducing the Galilei relativity

Let us modify condition (44) writing it in the form

$$\tau = \alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta} \cdot \vec{x}' .$$
<sup>(59)</sup>

where we have introduced the absolute time  $\tau$ . Therefore, assuming the Galilei transformation

$$\vec{x}' = \vec{x} - \vec{v}\tau$$

and using (52) and (59) we obtain that

$$\vec{v} = -\frac{2}{1 - \frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha^2} \,. \tag{60}$$

From (59) we also obtain

$$\frac{dt}{d\tau} = \frac{1}{\alpha c + \vec{\beta} \cdot \frac{d\vec{x}}{dt}} \; .$$

Let us define

$$\vec{u} := \frac{d\vec{x}}{d\tau} , \quad \vec{\tilde{u}} := \frac{d\vec{x}}{dt}$$
$$\vec{u}' := \frac{d\vec{x}'}{d\tau} , \quad \vec{\tilde{u}}' := \frac{d\vec{x}'}{dt}$$

then we obtain

$$\vec{u}\,' = \frac{\vec{\tilde{u}}\,'}{\alpha c - \vec{\beta} \cdot \vec{\tilde{u}}\,'} \,, \quad \vec{u} = \frac{\vec{\tilde{u}}}{\alpha c + \vec{\beta} \cdot \vec{\tilde{u}}} \,.$$

From the Galilei relativity we have

$$\vec{u}' = \vec{u} - \vec{v} \; ,$$

therefore, assuming  $\vec{u}' = 0$  we get  $\vec{u} = \vec{v}, \ \vec{\tilde{u}} = \vec{\tilde{v}}$  and

$$\vec{v} = \frac{\vec{\tilde{v}}}{\alpha c + \vec{\beta} \cdot \vec{\tilde{v}}} \; .$$

Replacing this last expression into (60) we end up with

$$\vec{\tilde{v}} = -\frac{2c}{1+\frac{\beta^2}{\alpha^2}}\frac{\vec{\beta}}{\alpha} \,. \tag{61}$$

In particular, we notice that

$$\widetilde{v} < c . \tag{62}$$

Then, the assumption of an absolute time and the Galilean relativity law in model II also leads to a situation where there are no superluminal frames.

The expression we obtained for  $\tilde{v}$  in (61) also leads to the same transformation of section 4.3.1, therefore, when we incorporate the Galilei relativity to the model II we reduce the model to the standard Lorentz situation. However, differently from the model I, as we have seen in section 4.3.1 we are allowed to obtain the Lorentz case without the need to impose the Galilei relativity.

#### 4.5 The momentum

We now search for an expression for the momentum that is conserved in collision processes. As it is well-known [5], the conservation of momentum in two frames S and S' brings to the analysis the transformation between them. Then, the problem here is to find an expression for the momentum that satisfies a conservation law as seen by frames related by the GLT.

Let us perform our analysis considering a particular collision process involving two particles with the same mass and such that relative to S we have before the collision the particles' velocities  $\vec{\tilde{u}}_1 = \tilde{u}\,\hat{\eta}, \,\,\vec{\tilde{u}}_2 = -\tilde{u}\,\hat{\eta}$ , and after the collision  $\vec{\tilde{u}}_1^* = \tilde{u}\,\hat{\eta}^*, \,\,\vec{\tilde{u}}_2^* = -\tilde{u}\,\hat{\eta}^*$ <sup>7</sup>. According to

<sup>&</sup>lt;sup>7</sup>As it is typical in the analysis of collision processes, for a more general collision we would obtain constraints on the final velocities, which are not essential in order to fixing the form of the function  $f(\vec{\tilde{u}})$  introduced in equations (63, 64).

S', we obtain from (56)

$$\begin{split} \vec{\tilde{u}}_{1}' &= \frac{\frac{2c}{1-\frac{\beta^{2}}{\alpha^{2}}} \frac{\vec{\beta}}{\alpha} + \widetilde{u} \, \vec{\omega}}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} + \widetilde{u} \, \frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})} \frac{\vec{\beta}}{\alpha} \cdot \widehat{\eta}} , \quad \vec{\tilde{u}}_{2}' &= \frac{\frac{2c}{1-\frac{\beta^{2}}{\alpha^{2}}} \frac{\vec{\beta}}{\alpha} - \widetilde{u} \, \vec{\omega}}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} - \widetilde{u} \, \frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})} \frac{\vec{\beta}}{\alpha} \cdot \widehat{\eta}} \\ \vec{\tilde{u}}_{1}'^{*} &= \frac{\frac{2c}{1-\frac{\beta^{2}}{\alpha^{2}}} \frac{\vec{\beta}}{\alpha} + \widetilde{u} \, \vec{\omega}^{*}}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} + \widetilde{u} \, \frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})} \frac{\vec{\beta}}{\alpha} \cdot \widehat{\eta}^{*}} , \quad \vec{\tilde{u}}_{2}' &= \frac{\frac{2c}{1-\frac{\beta^{2}}{\alpha^{2}}} \frac{\vec{\beta}}{\alpha} - \widetilde{u} \, \vec{\omega}^{*}}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} - \widetilde{u} \, \frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})} \frac{\vec{\beta}}{\alpha} \cdot \widehat{\eta}^{*}} \end{split}$$

with

$$\vec{\omega} := \hat{\eta} + \frac{2}{1 - \frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} \cdot \hat{\eta} \frac{\vec{\beta}}{\alpha} , \qquad \vec{\omega}^* := \hat{\eta}^* + \frac{2}{1 - \frac{\beta^2}{\alpha^2}} \frac{\vec{\beta}}{\alpha} \cdot \hat{\eta}^* \frac{\vec{\beta}}{\alpha}$$

being two vectors having arbitrary directions due to its dependence to the arbitrary unit vectors  $\hat{\eta}, \hat{\eta}^*$ . Following [5], we write for the momentum conservation in the frame S

$$f(\vec{\tilde{u}}_1)\vec{\tilde{u}}_1 + f(\vec{\tilde{u}}_2)\vec{\tilde{u}}_2 = f(\vec{\tilde{u}}_1^*)\vec{\tilde{u}}_1^* + f(\vec{\tilde{u}}_2^*)\vec{\tilde{u}}_2^*$$
(63)

and relative to S' it becomes (assuming the form of the function f is invariant)

$$f(\vec{\tilde{u}}_{1}')\vec{\tilde{u}}_{1}' + f(\vec{\tilde{u}}_{2}')\vec{\tilde{u}}_{2}' = f(\vec{\tilde{u}}_{1}'^{*})\vec{\tilde{u}}_{1}'^{*} + f(\vec{\tilde{u}}_{2}'^{*})\vec{\tilde{u}}_{2}'^{*}$$
(64)

that is equivalent to

$$\begin{split} \widetilde{u}\,\widetilde{\omega}\,\left\{\frac{f(\vec{\tilde{u}}_{1}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}+\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}}-\frac{f(\vec{\tilde{u}}_{2}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}-\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}}\right\}+\\ -\widetilde{u}\,\vec{\omega}^{*}\,\left\{\frac{f(\vec{\tilde{u}}_{1}'^{*})}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}+\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}}-\frac{f(\vec{\tilde{u}}_{2}'^{*})}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}-\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}}\right\}+\\ +\frac{\vec{\beta}}{\alpha}\,\frac{2c}{1-\frac{\beta^{2}}{\alpha^{2}}}\,\left\{\frac{f(\vec{\tilde{u}}_{1}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}+\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}}+\frac{f(\vec{\tilde{u}}_{2}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}-\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}}+\\ -\frac{f(\vec{\tilde{u}}_{1}'^{*})}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}+\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}}-\frac{f(\vec{\tilde{u}}_{2}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}}-\widetilde{u}\,\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}}\right\}=0\,. \end{split}$$

From the arbitrariness of the scattering directions  $\widehat{\eta},\,\widehat{\eta}^{\,*}$  we must have

$$\frac{f(\vec{\tilde{u}}_{1}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} + \widetilde{u}\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}} = \frac{f(\vec{\tilde{u}}_{2}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} - \widetilde{u}\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}} = \frac{f(\vec{\tilde{u}}_{1}'^{*})}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} - \widetilde{u}\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}} = \frac{f(\vec{\tilde{u}}_{2}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} - \widetilde{u}\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}} = \frac{f(\vec{\tilde{u}}_{2}')}{\frac{1+\frac{\beta^{2}}{\alpha^{2}}}{1-\frac{\beta^{2}}{\alpha^{2}}} - \widetilde{u}\frac{2}{c(1-\frac{\beta^{2}}{\alpha^{2}})}\frac{\vec{\beta}}{\alpha}\cdot\widehat{\eta}^{*}}.$$
(65)

In order to satisfy this condition we notice that from (57) we obtain

$$\sqrt{\left|1 - \frac{\tilde{u}'^2}{c^2}\right|} = \frac{\sqrt{\left|1 - \frac{\tilde{u}^2}{c^2}\right|}}{\left|\frac{1 + \frac{\beta^2}{\alpha^2}}{1 - \frac{\beta^2}{\alpha^2}} + \frac{2}{1 - \frac{\beta^2}{\alpha^2}}\frac{\vec{\beta}}{\alpha} \cdot \frac{\vec{u}}{c}\right|}.$$
(66)

For the both cases we analyzed in sections 4.3.1, 4.3.2, we observe that  $\frac{\beta}{\alpha} < 1$  and for the case of a subluminal particle, i.e. u < c, u' < c, we obtain

$$\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}} + \frac{2}{1-\frac{\beta^2}{\alpha^2}}\frac{\vec{\beta}}{\alpha} \cdot \frac{\vec{\tilde{u}}}{c} > 0$$
(67)

which suggest us to take for  $f(\tilde{u})$  the following expression

$$f(\widetilde{u}) = \frac{k}{\sqrt{1 - \frac{\widetilde{u}^2}{c^2}}}$$

In the case of a superluminal particle, i.e. u > c, u' > c, we have in general

$$\frac{1+\frac{\beta^2}{\alpha^2}}{1-\frac{\beta^2}{\alpha^2}} + \frac{2}{1-\frac{\beta^2}{\alpha^2}}\frac{\vec{\beta}}{\alpha} \cdot \frac{\vec{\tilde{u}}}{c} < 0$$
(68)

which suggest us to take

$$f(\widetilde{u}) = \frac{k \operatorname{sgn} \left[ 1 + \frac{\beta^2}{\alpha^2} + 2\frac{\vec{\beta}}{\alpha} \cdot \frac{\vec{\tilde{u}}}{c} \right]}{\sqrt{\left| 1 - \frac{\widetilde{u}^2}{c^2} \right|}} .$$
(69)

The constant k is identified with the rest mass of the particle in the case one can effectively define such frame. In the case of tachyons we refer to the discussion of [13]. The term in the numerator,  $\operatorname{sgn}\left[1 + \frac{\beta^2}{\alpha^2} + 2\frac{\vec{\beta}}{\alpha} \cdot \frac{\vec{u}}{c}\right]$ , becomes necessary in order to fulfill the conservation of momentum as expressed in condition (65). This term introduces a dependence of the tachyon mass with the quantity  $\frac{\vec{\beta}}{\alpha}$ , which is ultimately a parameter depending on the relative velocity between the frames. A detailed discussion on the implications of this term may be found in [13].

# 5 An alternative way to induce a concept of absolute time

The concept of the absolute time was introduced in models I and II by means of certain relations involving the space and time variables, equations (4), (59). These relations may be seen as kinematical requirements allowing to bring the special and the Galilei relativity into the models. In this view, we could think on the physical and the absolute time without having to consider them as coordinates of a 4-dimensional space. Now, we reverse the construction and we will show that by considering a higher dimensional space we can think on the absolute time with no reference to kinematical relations.

### 5.1 Setting the Lorentz transformation in $\mathbb{R}^{(1,4)}$

We denote by  $\mathbb{R}^{(1,4)}$  the space  $\mathbb{R}^5$  endowed with the quadratic form

$$Q(x^{a}) = -(x^{0})^{2} + (x^{1})^{2} + (x^{2})^{2} + (x^{3})^{2} + (x^{4})^{2}.$$

Here  $\mathbb{R}^{(1,4)} \ni x \equiv (x^a), a = 0, 1, 2, 3, 4$ . The transformations leaving Q invariant have the form

$$x^{\prime a} := \Lambda^a{}_b x^b \tag{70}$$

with  $[\Lambda^a_b] \in SO(1,4)$ . We consider now two subgroups of SO(1,4) as follows. The first consists on transformations defined by matrices

$$L_{\Lambda} := \left[ \begin{array}{cc} \Lambda^{\mu}_{\ \nu} & \mathbf{0} \\ \mathbf{0} & 1 \end{array} \right]$$

with  $\Lambda \in SO(1,3)$  ( $\mu, \nu = 0, 1, 2, 3$ ). The second consists on transformations defined by matrices

$$L_{\mathcal{R}} := \left[ egin{array}{cc} 1 & \mathbf{0} \\ \mathbf{0} & \mathcal{R}^i_j \end{array} 
ight]$$

with  $\mathcal{R} \in SO(4)$  (i, j = 1, 2, 3, 4).

Now, we define new coordinates as follows:

(i)  $(x^a/x^4) = (\hat{x}^{\mu}, 1)$  with  $\hat{x}^{\mu} \equiv x^{\mu}/x^4$ . For the subgroup  $\{L_{\Lambda}\} \subset SO(1, 4)$  the transformation (70) becomes

$$x'^{\mu} = \Lambda^{\mu}_{\ \nu} x^{\nu} \qquad x'^4 = x^4 \tag{71}$$

and in particular

 $\widehat{x}^{\prime\mu} = \Lambda^{\mu}_{\ \nu} \widehat{x}^{\nu} \; .$ 

(ii)  $(x^a/x^0) = (1, \tilde{x}^i)$ , with  $\tilde{x}^i \equiv x^i/x^0$ . For the subgroup  $\{L_{\mathcal{R}}\} \subset SO(1, 4)$  the transformation (70) becomes

$$x^{\prime 0} = x^0 \qquad x^{\prime i} = \mathcal{R}^i_i x^j \tag{72}$$

and in particular

$$\widetilde{x}^{\prime i} = \mathcal{R}^i_j \widetilde{x}^j$$

We notice that these coordinates  $\hat{x} \equiv (\hat{x}^{\mu})$  and  $\tilde{x} \equiv (\tilde{x}^{i})$  can't be identified with the coordinates of the hyperplanes  $x^{0} = 0$ ,  $x^{4} = 0$  of  $\mathbb{R}^{(1,4)}$ .

## 5.2 Inducing Lorentz transformation in Euclidean space $\mathbb{R}^4$

There is a natural way to reinterpret the coordinates  $\hat{x} = (\hat{x}^{\mu}), \, \tilde{x} = (\tilde{x}^{i})$  as local coordinates on certain open sets of the real projective space  $\mathbb{RP}^{4}$  [14]. In fact, let  $\mathbb{RP}^{4} \ni [x^{0}, x^{1}, x^{2}, x^{3}, x^{4}] :=$ 

 $\{\lambda(x^0, x^1, x^2, x^3, x^4) : \lambda \in \mathbb{R}^*\}$ . Recall that a smooth atlas for  $\mathbb{RP}^4$  is given by  $\mathbb{RP}^4 := \{(V_a, \varphi_a)\}_{a=0,\dots,4}$  where  $V_a := \{[x^0, x^1, x^2, x^3, x^4] : x^a \neq 0\}$  and

$$\varphi_a : V_a \to \mathbb{R}^4$$
$$[x] \equiv [x^0, x^1, x^2, x^3, x^4] \to \varphi_a([x^0, x^1, x^2, x^3, x^4]) := \left(\frac{x^0}{x^a}, \cdots, \frac{x^{a-1}}{x^a}, \frac{x^{a+1}}{x^a}, \cdots, \frac{x^4}{x^a}\right)$$

which shows that  $\hat{x} = \varphi_4([x])$ , and  $\tilde{x} = \varphi_0([x])^{-8}$ .

The transformation on  $\mathbb{R}^{(1,4)}$ ,  $x'^a = \Lambda^a_b x^b$ , induces a transformation on  $\mathbb{R}P^4 : [x] \to [\Lambda x]$  that has the particular cases

$$\begin{split} & [x^0, x^1, x^2, x^3, x^4] \xrightarrow{L_{\Lambda}} [\Lambda^0_{\ \alpha} x^{\alpha}, \Lambda^1_{\ \alpha} x^{\alpha}, \Lambda^2_{\ \alpha} x^{\alpha}, \Lambda^3_{\ \alpha} x^{\alpha}, x^4] \\ & [x^0, x^1, x^2, x^3, x^4] \xrightarrow{L_{\mathcal{R}}} [x^0, \mathcal{R}^1_i x^i, \mathcal{R}^2_i x^i, \mathcal{R}^3_i x^i, \mathcal{R}^4_i x^i] \; . \end{split}$$

Therefore,  $\varphi_4 \circ L_\Lambda \circ \varphi_4^{-1}$  reduces to an ordinary Lorentz transformation in  $\varphi_4(V_4) \simeq \mathbb{R}^4$ :  $\hat{x}^{\mu} \longrightarrow \hat{x}'^{\mu} = \Lambda^{\mu}_{\nu} \hat{x}^{\nu}$ , while  $\varphi_0 \circ L_{\mathcal{R}} \circ \varphi_0^{-1}$  reduces to a 4-dimensional rotation in  $\varphi_0(V_0) \simeq \mathbb{R}^4$ :  $\tilde{x}^i \longrightarrow \tilde{x}'^i = \mathcal{R}^i_j \tilde{x}^j$ . We have then modelled Minkowski and Euclidean four-space as homeomorphic to open sets  $\varphi_4(V_4)$ ,  $\varphi_0(V_0)$  of projective space where Lorentz and 4-dimensional rotations are induced by transformations (71, 72) of SO(1, 4) acting in  $\mathbb{R}^5$ . In the intersection of the neighborhoods,  $V_0 \cap V_4$ , we have  $\tilde{x}^i = \varphi_0 \circ \varphi_4^{-1}(\hat{x})$ , i.e.

$$\widetilde{x}^1 = \frac{\widehat{x}^1}{\widehat{x}^0}, \quad \widetilde{x}^2 = \frac{\widehat{x}^2}{\widehat{x}^0}, \quad \widetilde{x}^3 = \frac{\widehat{x}^3}{\widehat{x}^0}, \quad \widetilde{x}^4 = \frac{1}{\widehat{x}^0}.$$
(73)

Therefore, in the intersection  $V_0 \cap V_4$  an ordinary Lorentz transformation on  $\varphi_4(V_4)$ ,  $\hat{x}^{\mu} \to \hat{x}'^{\mu} = \Lambda^{\mu}_{\nu} \hat{x}^{\nu}$ , induces on  $\varphi_0(V_0)$  the transformation

$$\widetilde{x}^{i} \to \widetilde{x}^{\prime i} = \frac{L^{i}_{\Lambda 0} + L^{i}_{\Lambda j} \widetilde{x}^{j}}{L^{0}_{\Lambda 0} + L^{0}_{\Lambda j} \widetilde{x}^{j}}$$
(74)

that also obeys a group law, e.g.

i.e.

$$\widetilde{x}_{(2)}^{i} = \frac{(L_{(2)}L_{(1)})^{i}{}_{0} + (L_{(2)}L_{(1)})^{i}{}_{j}\widetilde{x}^{j}}{(L_{(2)}L_{(1)})^{0}{}_{0} + (L_{(2)}L_{(1)})^{0}{}_{j}\widetilde{x}^{j}} \ .$$

Equation (74) may be seen as a non-linear realization of the Lorentz group in  $\varphi_0(V_0)$ .

<sup>&</sup>lt;sup>8</sup>Notice that the space  $\mathbb{R}^4$ , appearing as the image of the homeomorphisms  $\varphi_a$ , is not seen as a subset of  $\mathbb{R}^{(1,4)}$ .

# 5.3 Interpreting the boost transformation in Euclidean space $\mathbb{R}^4$

Now, let us introduce some physics in our analysis. So far, we have considered homogeneous coordinates  $\hat{x}, \tilde{x}$  that carry no dimension. In order to give a physical meaning to them let us assume there is a scale factor  $\lambda$  with dimension of length and let us define  $y^i := \lambda \tilde{x}^i, y'^i := \lambda' \tilde{x}'^i$ . Here  $\lambda'$  is the transformed scale factor whose form is left arbitrary for a moment. Then, transformation (74) may be seen as a transformation

$$y^{i} \rightarrow y^{\prime i} = \lambda^{\prime} \frac{L^{i}_{\Lambda 0} \lambda + L^{i}_{\Lambda j} y^{j}}{L^{0}_{\Lambda 0} \lambda + L^{0}_{\Lambda j} y^{j}} .$$

$$\tag{75}$$

For a boost we have  $\Lambda_0^0 = \gamma$ ,  $\Lambda_{\overline{i}}^0 = \Lambda_{\overline{i}}^{\overline{i}} = -\frac{\gamma}{c}v^{\overline{i}}$ ,  $\Lambda_{\overline{j}}^{\overline{i}} = \delta_{\overline{j}}^{\overline{i}} - (1-\gamma)\frac{v^{\overline{i}}v^{\overline{j}}}{v^2}$ , and we obtain

$$y^{\prime \bar{i}} = \lambda^{\prime} \frac{y^{\bar{i}} - (1 - \gamma) \frac{y^{\bar{j}} v^{\bar{j}}}{v^2} v^{\bar{i}} - \frac{\gamma}{c} \lambda v^{\bar{i}}}{\gamma(\lambda - \frac{1}{c} y^{\bar{j}} v^{\bar{j}})}$$

$$y^{\prime 4} = \lambda^{\prime} \frac{y^4}{\gamma(\lambda - \frac{1}{c} y^{\bar{j}} v^{\bar{j}})}$$
(76)

where  $\bar{i} = 1, 2, 3$ . Now, let us assume the scale factor writes as  $\lambda' = ct', \lambda = ct$  with c the speed of light and t and t' having the dimension of time. Here, denoting  $\vec{y} = (y^{\bar{i}}), \vec{v} = (v^{\bar{i}})$  equation (76) becomes

$$\vec{y}' = t' \frac{\vec{y} - (1 - \gamma) \frac{\vec{y} \cdot \vec{v}}{v^2} \vec{v} - \gamma t \vec{v}}{\gamma (t - \frac{\vec{y} \cdot \vec{v}}{c^2})}$$

$$y'^4 = t' \frac{y^4}{\gamma (t - \frac{\vec{y} \cdot \vec{v}}{c^2})}$$

If  $t' = \gamma(t - \frac{\vec{y} \cdot \vec{v}}{c^2})$  we obtain

$$\vec{y}' = \vec{y} - (1 - \gamma) \frac{\vec{y} \cdot \vec{v}}{v^2} \vec{v} - \gamma t \vec{v}$$

$$y'^4 = y^4.$$
(77)

Eq. (77) together with  $t' = \gamma(t - \frac{\vec{y} \cdot \vec{v}}{c^2})$  constitute the standard transformation of special relativity, and  $y'^4 = y^4$  can be considered as defining the absolute time, e.g.  $\tau \equiv y'^4/c = y^4/c$ . In this way, the absolute time acquires an intrinsic characteristic as the coordinate (in fact with  $\frac{y^4}{c}$ ) of a 4-dimensional Euclidean space  $\mathbb{R}^4 \simeq \varphi_0(V_0)$ , while the physical time of SR appears as a non-invariant scale factor (through the relation  $\lambda = ct, \lambda' = ct'$ ) transforming homogeneous coordinates  $\tilde{x}^i$  of Euclidean space  $\mathbb{R}^4$  into dimensional coordinates  $y^i$ . If in addition we impose the kinematical relation of the Galilei relativity,  $\vec{y}' = \vec{y} - \vec{v}\tau$ , we end up with  $\tau = (1-\gamma)\frac{\vec{y}\cdot\vec{v}}{v^2}\vec{v} + \gamma\vec{v}t$ .

# 6 Conclusion

In our work we have consistently unified the Galilei and the Special relativity. In Model I the main equations were  $\vec{x}' = \vec{x} - \vec{v}$  and  $\vec{x}'^2 - c^2 t'^2 = \vec{x}^2 - c^2 t^2$ , while in Model II we have used

 $\vec{x}' = \vec{x} - \vec{v}\tau$  and  $\tau = \alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta} \cdot \vec{x}'$ . Here, contrarily to the intrinsic geometric meaning associated to the relation  $\vec{x}'^2 - c^2 t'^2 = \vec{x}^2 - c^2 t^2$ , it is not clear the role played by  $\alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta} \cdot \vec{x}'$ . If there is a certain anisotropy in space, represented by a particular direction in space (not necessarily related to the relative velocity of the two frames), then we could encode it into  $\vec{\beta}$ . In this way, the physics derived from the general transformation (52) may provide a framework to describe situations other than the one we contemplated in our work.

The assumption of the absolute time leads to a distinction between the velocities  $\vec{v} = \frac{d\vec{r}}{d\tau}$ and  $\vec{v} = \frac{d\vec{r}}{dt}$ . As we have seen in (19, 62), the Galilei relativity imposes a relation between these velocities in such way that  $\tilde{v}$  becomes less than c, which rules out tachyons since in SR we consider velocities only as derivatives with respect to t. However, as we have seen in Model II, it is possible to have superluminal frames provided we don't impose the existence of the absolute time and the Galilei relativity. Having distinguished between the absolute and the physical time, the assumption of the constancy of the speed of light is understood only with respect to the derivative relative to the physical time. In fact, as we have seen in the discussion of section 3.4.2, for the light speed calculated as a derivative relative to the absolute time we obtained  $c_S \neq c_{S'}$ .

Finally, the implications of the existence of an absolute time in the formulation of a quantum theory is an old topic. Dirac, for example, has used it in his attempt to develop electrodynamics [15]. In the context of our work, one possibility that arises is related to the development of quantum theory following the Schwinger Quantum Action Principle (SQAP). When applied to quantum mechanics this approach uses an hamiltonian formalism that singles out the time as parameter. As we have shown in [16], a consequence of the application of a modified form of the SQAP to quantum mechanics has lead to the commutator of the Lie algebra of the Galilei group. Now, we could apply the same procedure using the absolute time as parameter together with its relation to the physical time as given by (44) and search for the corresponding modifications it would originate in the Galilei algebra. Then, we could check if it corresponds to the group of transformation leaving the quantity  $\tau = \alpha x^0 + \vec{\beta} \cdot \vec{x} = \alpha x'^0 - \vec{\beta} \cdot \vec{x}'$  invariant.

Acknowledgements: We thank all the staff of the library of Centro Brasileiro de Pesquisas Físicas (CBPF), in particular to Vera, Fátima, Edileuza, Rosa and Ramalho. M.C. thanks Aurelina Carvalho, José E. Carvalho, Aureliana Cabral Raposo and the enlightened conversations with Yohan Ricardo. A.L. thanks Eliana N. Lyra de Oliveira for the permanent support and Maria I. Lyra (*in memoriam*), and dedicates this work (*in memoriam*) to his sister Anita Lyra de Oliveira with Down syndrome, recently deceased. Both authors thank prof. A. F. da F. Teixeira of CBPF for some comments on part of our work.

In honour of  $\overline{\text{IC}} \overline{\text{XC}}$ ,  $\overline{\text{MP}} \overline{\Theta \Upsilon}$ .

# References

[1] Sela, O., Tamir, B., Dolev, S., Elitzur, A.: Can Special Relativity Be Derived from Galilean

Mechanics Alone?, Found. Phys. 39 (2009) 499.

- [2] Cahill, R. T.: Unravelling Lorentz Covariance and the Spacetime Formalism, Prog. Phys. 4 (2008) 19.
- [3] Desloge, E.A, Philpott, R.J.: Uniformly accelerated reference frames in special relativity, Am. J. Phys. 55 (1987) 252.
- [4] French, A. P.: Newtonian Mechanics, W.W. Norton and Co. (1971), N.Y.
- [5] Moller, C.: Theory of Relativity, Oxford Univ. Press (1972), Oxford.
- [6] Desloge, E.A.: Classical Mechanics, Vol. 1, Wiley (1986), N.Y.
- [7] Laudau, L., Lifshitz, E.: The Classical Theory of Fields, Butterworth-Heinemann (1980).
- [8] Einstein, A.: *Relativity*, Three Rivers Press (1961) NY.
- [9] Shankara, T. S.: Tachyons via Supersonics, Found. Phys. 4 (1974) 97.
- [10] Duffey, G. H.: Tachyons and Superluminal Wave Groups, Found. Phys. 5 (1975) 349.
- [11] Ritus, V. I.: On the difference between Wigner's and Moller's approaches to the description of Thomas precession, Physics - Uspekhi 50(2007) 95.
- [12] Kapuscik, E.: Special Theory of Relativity without special assumptions and tachyonic motion, Cond. Matter Phys. 13, No 4, (2010) 43102:1 – 8.
- [13] Park, Mu-In, Park, Young-Jai: On the foundation of the relativistic dynamics with the tachyon, Il Nuovo Cimento 111B (1996) 1333.
- [14] do Carmo, M. P.: Geometria Riemaniana. Projeto Euclides. Impa, Rio de Janeiro (1988)
- [15] Dirac, P.A.M.: The Lorentz Transformation and Absolute Time, Physica XIX (1953) 888.
- [16] Carvalho, M., Lyra, A.: The Quantum Action Principle Revisited, Canadian Journ. of Phys. 86 (2008) 703.