

1 A Method for Detecting Lagrangian Coherent
2 Structures (LCSs) using Fixed Wing
3 Unmanned Aircraft System (UAS)

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Abstract

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TBD

15 **1 Introduction**

16 The transport of material in the atmosphere is a problem with important im-
17 plications for agriculture [1–4], aviation [5, 6], and human health [7, 8]. Given
18 the turbulent nature of the atmosphere it can be difficult to predict where
19 a particle, such as a plant pathogen, will wind up. Tools from dynamical
20 systems theory, such as LCSs, can help us to understand how particles in a
21 flow will evolve. The study of transport in the atmosphere from a dynamical
22 systems perspective has long focused on the study of large scale phenom-
23 ena [1–5, 9–12]. This has been largely due to the larger scale grid spacing
24 of readily available atmospheric model data and the lack of high resolution
25 atmospheric measurements on a scale large enough to calculate Lagrangian
26 data. Furthermore, few works have attempted to find ways to detect LCSs

27 in the field. In [6, 13] the authors used wind velocity measurements from a
28 dopler LiDAR to detect LCS which had passed through Hong Kong Inter-
29 national Airport. The authors in [1] took a different, rather than measure
30 the wind velocity to try and detect LCSs, the authors looked at sudden
31 changes in pathogen concentrations in the atmosphere and were able to link
32 those changes to the passage of LCSs using atmospheric forecasts from the
33 North American Mesoscale (NAM) model. Yet to date, we are unaware of
34 any attempts to develop a means of directly sense LCSs in which could be
35 readily implemented by operators in the field. Recent advances in dynamical
36 systems theory, such as new Eulerian diagnostics, as well as, atmospheric
37 sensing technology, such as unmanned aircraft systems (UAS), have brought
38 the detection of localized LCSs within reach.

39 The first of these developments are new Eulerian techniques for measuring
40 the attraction and repulsion of regions in a fluid flow [14, 15]. In traditional
41 Lagrangian analyses a velocity field is needed which is defined over a large
42 enough spatiotemporal scale for the advection of virtual particles. These new
43 Eulerian methods do not rely on the advection of virtual particles, instead
44 they utilize the gradients of the velocity field. Since they rely on gradients,
45 these techniques we only require enough points to enact a finite-differencing
46 scheme. Furthermore, these methods are Eulerian and thus can be made
47 using temporally coarse or even temporally pointwise data sets.

48 The second of these developments is the use of inexpensive UAS to sam-
49 ple the atmospheric velocity instead of piloted aircraft or other traditional
50 assets. Ground-based wind sensors such as LiDAR (light detection and rang-
51 ing), SoDAR (sonic detection and ranging), or tower-mounted anemometers
52 can be prohibitively expensive and difficult to relocate to regions of interest,
53 such as a hazardous zone. Airborne wind measurement from aircraft has a
54 long history [16, 17] and well-developed existing programs [18]. The prolif-
55 eration of unmanned aircraft systems (UAS) has enabled wind measurement
56 missions which may be lower cost, longer duration, or in more dangerous
57 environments. Elston et. al. [19] provide a review of many UAS atmospheric
58 measurement efforts, and recent works continue to advance both theoretical
59 and practical UAS capabilities [20–25].

60 In this paper we will take advantage of these developments to advance a
61 methodology, based on numerical simulations, which will enable UAS oper-
62 ators in the field to utilize their wind measurements to detect LCSs.

63 2 Methods

64 2.1 Lagrangian-Eulerian Analysis

65 We will be analyzing the dynamical system

$$\frac{d}{dt}\mathbf{x}(t) = \mathbf{v}(\mathbf{x}(t), t), \quad (1)$$

$$\mathbf{x}_0 = \mathbf{x}(t_0). \quad (2)$$

66 In this system $\mathbf{x}(t)$ is the position vector of a fluid parcel at time t and
67 $\mathbf{v}(\mathbf{x}, t)$ is the horizontal wind velocity vector at position $\mathbf{x}(t)$, time t . We
68 define the components of the horizontal position vector, $\mathbf{x} = (x, y)$, where x
69 is the eastward position and y is the northward position and the horizontal
70 velocity vector, $\mathbf{v} = (u, v)$, where u is the eastward velocity and v is the
71 northward velocity.

72 We will be analyzing this system using both Lagrangian and Eulerian
73 tools. For the Lagrangian analysis we will be using the Finite-Time Lyapunov
74 Exponent (FTLE), σ , and Lagrangian coherent structures (LCSs). We de-
75 fine LCSs as C-ridges of the FTLE field following [26]. The FTLE field is a
76 measure of the stretching of a fluid parcels within a flow, the forward-time
77 FTLE measures repulsion and the backward-time FTLE measures attrac-
78 tion. LCSs on the other hand are the most attracting and repelling material
79 surfaces within a fluid flow; they provide a means of visualizing how particles
80 within the flow will evolve.

81 For the Eulerian analyses we use the attraction rate, s_1 , and the trajectory
82 divergence rate, ρ , both of which are derived from the Eulerian rate-of-strain
83 tensor, \mathbf{S} , described in equation 6. The attraction rate is the minimum eigen-
84 value of \mathbf{S} and was shown in ref [14] to provide a measure of instantaneous
85 hyperbolic attraction, with isolated minima of s_1 providing the cores of at-
86 tracting objective Eulerian coherent structures (OECS). Recent work has
87 shown that in 2D, s_1 is the limit of the backward-time FTLE as integration
88 time goes to 0 [27]. The trajectory divergence rate is a measure of how the
89 how much repulsion is changing along streamlines of the velocity field.

90 To calculate Lagrangian metrics we must first calculate the flow map for
91 the time period of interest,

$$\mathbf{F}_{t_0}^t(\mathbf{x}_0) = \mathbf{x}_0 + \int_{t_0}^t \mathbf{v}(\mathbf{x}(t), t) dt. \quad (3)$$

92 Taking the gradient of the flow map we can then calculate the right Cauchy-
 93 Green strain tensor,

$$\mathbf{C}_{t_0}^t(\mathbf{x}_0) = \nabla \mathbf{F}_{t_0}^t(\mathbf{x}_0)^T \cdot \nabla \mathbf{F}_{t_0}^t(\mathbf{x}_0), \quad (4)$$

94 From the largest eigenvalue of the right Cauchy-Green strain tensor, λ_n ,
 95 we can then calculate the FTLE field,

$$\sigma_{t_0}^t(\mathbf{x}_0) = \frac{1}{2|t - t_0|} \log(\lambda_n(\mathbf{x}_0)) \quad (5)$$

96 For the Eulerian metrics, the Eulerian rate-of-strain tensor is defined as

$$\mathbf{S}(\mathbf{x}_0) = \frac{1}{2} (\nabla \mathbf{v}(\mathbf{x}_0) + \nabla \mathbf{v}(\mathbf{x}_0)^T). \quad (6)$$

97 The attraction rate, s_1 , is the minimum eigenvalue of \mathbf{S} . The trajectory
 98 divergence rate is defined as

$$\dot{\rho}(\mathbf{x}_0) = \hat{\mathbf{n}}(\mathbf{x}_0)^T \cdot \mathbf{S}(\mathbf{x}_0) \cdot \hat{\mathbf{n}}(\mathbf{x}_0) = \frac{1}{\|\mathbf{v}(\mathbf{x}_0)\|^2} (\mathbf{v}(\mathbf{x}_0)^T \cdot \mathbf{J}^T \cdot \mathbf{S}(\mathbf{x}_0) \cdot \mathbf{J} \cdot \mathbf{v}(\mathbf{x}_0)), \quad (7)$$

99 where $\hat{\mathbf{n}}(\mathbf{x}_0)$ is the unit vector normal to the trajectory and \mathbf{J} is the symplectic
 100 matrix [15]. A visual interpretation of the trajectory divergence rate can be
 101 found in figure 1.

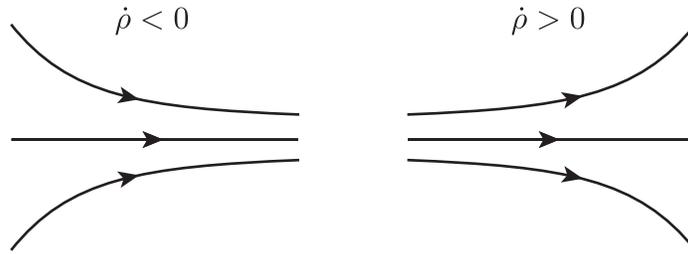


Figure 1: Schematic of the trajectory divergence rate, taken from [15]. Where $\dot{\rho} < 0$, trajectories are converging, where $\dot{\rho} > 0$ trajectories are diverging.

102 **2.2 Gradient Approximation from UAS Flight Data**

103 In order to calculate the Eulerian rate-of-strain tensor from our UAS data
104 sets we have developed an algorithm to calculate the gradient of a scalar field
105 based on measurements along a circular arc. An assumption that goes into
106 this algorithm is that the scalar field is not significantly changing in time
107 during the period of one full orbit, but is changing in space. We believe this
108 assumption is appropriate to apply to atmospheric velocity fields, as mid to
109 larger scale atmospheric flows tend to change on the order of hours, while
110 UAS orbits are on the order of minutes. This assumption of course ignores
111 small scale turbulent motion which would fall below the scale at which we
112 are sampling, *m vs km*. This algorithm also assumes that the important
113 features will be in the horizontal plane. This assumption was previously
114 applied to atmospheric model data in [9, 10, 28] based on the fact that the
115 vertical component of the wind velocity tends to be two orders of magnitude
116 less than the horizontal components.

117 This algorithm takes the radius of the circle, r , which is assumed to be
118 constant, as a scalar input, the angle θ as an $n \times 1$ array input, and a scalar u
119 as an $n \times 1$ array input. Note, this algorithm is currently written for a clock-
120 wise trajectory, however, it would work equally well for a counterclockwise
121 trajectory with the appropriate modifications. We start with an initial point
122 along the circular flight path (r, θ_0) and u at that point, then provided the
123 path continues for at least another $\frac{3}{4}$ of a circle, interpolate u to 3 additional
124 points along the path at $(r, \theta_0 - \frac{1}{2}\pi)$, $(r, \theta_0 - \pi)$, and $(r, \theta_0 - \frac{4}{3}\pi)$. With u at
125 4 individual points along the flight path use a central difference scheme to
126 approximate the gradient of u at the center point of the circular path. Since
127 these 4 points are along an arc, the gradient of each set of 4 points will be
128 in a different frame of reference from our initial set. To correct for this we
129 apply a counterclockwise rotation to the gradient vector of u to obtain the
130 gradient in our reference frame. Continue this method for each additional
131 point along the circular path until there is less than an additional $\frac{3}{4}$ of a circle
132 left. A pseudo-code version of this algorithm can be found in Algorithm 1,
133 and a schematic can be found in figure 2.

134 **2.3 Model Data**

135 For a velocity field we used data from the 3km North American Mesoscale
136 (NAM) model. We looked at a section of the model over Southwestern Vir-

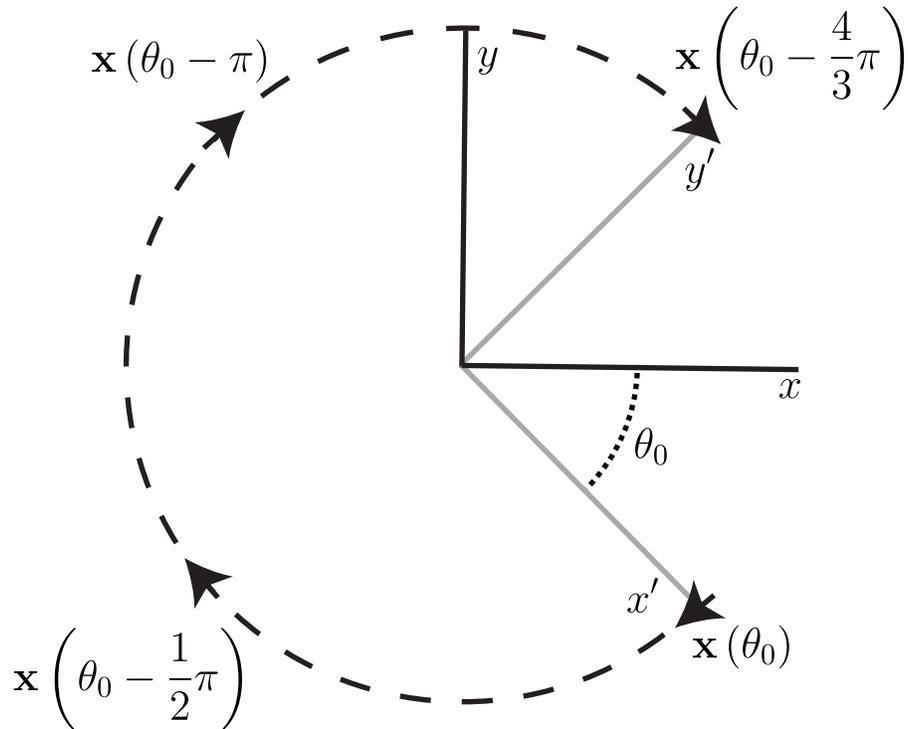


Figure 2: Schematic showing positions where velocity measurements were made and the position of the circle gradient frame to the reference frame.

137 ginia centered at the Virginia tech experimental farm during a 215hr period
 138 beginning Sept 4th, 2017 at 00:00 UTC. We divided the NAM data was into
 139 2 parts. The first part was a strictly 2D data set that looked at the 850mb
 140 isosurface, the second was a 3D data set. Both data sets were interpolated in
 141 time from 1hr resolution to 10min resolution using cubic splines. The 3D data
 142 was then interpolated from pressure based vertical levels to height based ver-
 143 tical levels using linear interpolation from MATLAB's scatteredInterpolant
 144 routine. Both data sets were also interpolated from a 3km horizontal resolu-
 145 tion to a 300m horizontal resolution using cubic Lagrange polynomials. The
 146 3D data set was then fed to a flight simulator which attempted the follow
 147 the 850mb isosurface. A subscale model of a transport-style aircraft, named
 148 the T-2, was used as the simulated unmanned aircraft. To get a sense of its
 149 scale, some of its common physical properties are

$$\text{mass } m = 22.5 \text{ kg} \quad \text{wingspan } b = 2.09 \text{ m} \quad \text{chord } \bar{c} = 0.28 \text{ m}$$

Algorithm 1 Circle Gradient

```
1: input  $\theta, u(\theta), r$ 
2: for  $i$  in length( $\theta$ ) do
3:   if  $\theta(i) - \frac{4}{3}\pi \geq \theta(\text{end})$  then
4:     interpolate  $u(\theta)$  to  $u(\theta(i) - \frac{1}{2}\pi), u(\theta(i) - \pi), u(\theta(i) - \frac{4}{3}\pi)$ 
5:      $\frac{du}{dx'} = \frac{u(\theta(i)) - u(\theta(i) - \pi)}{2 \cdot r}$ 
6:      $\frac{du}{dy'} = \frac{u(\theta(i) - \frac{1}{2}\pi) - u(\theta(i) - \frac{4}{3}\pi)}{2 \cdot r}$ 
7:      $\nabla u(i) = \begin{bmatrix} \cos(\theta(i)) & \sin(\theta(i)) \\ -\sin(\theta(i)) & \cos(\theta(i)) \end{bmatrix} \begin{bmatrix} \frac{du}{dx'} \\ \frac{du}{dy'} \end{bmatrix}$ 
return  $\nabla u$ 
```

150 The T-2 cruising airspeed is approximately 40 m/s. The details of the flight
151 dynamic model are included in Appendix A. The simulated wind “measure-
152 ments” taken by the aircraft are wind field components along the aircraft’s
153 center-of-mass trajectory $\mathbf{v}(\mathbf{x}(t), t)$.

154 3 Results and Discussion

155 3.1 Approximating local Eulerian Metrics from UAS 156 flights

157 In this section we examine how well the attraction rate, s_1 , and the trajec-
158 tory divergence rate, $\dot{\rho}$, can be approximated from a UAS flight. Figure 3
159 shows the results for the trajectory divergence rate. Using the 850mb isosur-
160 face velocity field we calculated the trajectory divergence rate at the center
161 point of our circle/flight radius, shown in red. We then used velocity data
162 from a perfectly circular path with a radius varying from 2km to 15km re-
163 stricted to the 850mb isosurface to approximate the trajectory divergence
164 rate, shown in black. Finally, we used velocity data from a 3D simulated
165 UAS flight path with a radius varying from 2km to 15km attempting to
166 follow the 850mb isosurface to approximate the trajectory divergence rate,
167 shown in blue. Pearson correlation coefficients for these measurements can
168 be found in table 3.1.

169 We can see from the results in figure 3 that the simulated UAS flight in a
170 3D space provides a very similar result to the circular path restricted to the
171 850mb isosurface. For all the radii we looked at the trajectory divergence

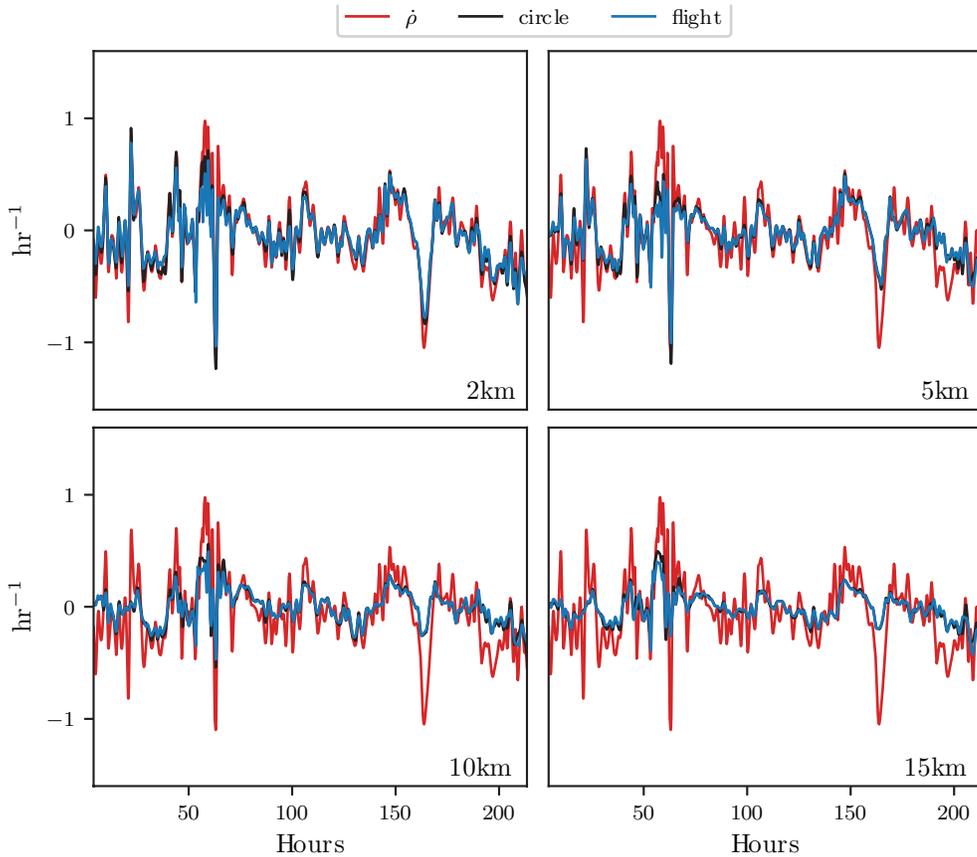


Figure 3: Comparison of the trajectory divergence rate measurements between center point (red), circular arc (black), and simulated drone flight (blue). Radius for the measurements is shown in lower right hand corner.

172 rate from the flight simulation is nearly identical to that from the 2D circular
 173 path. Most of the error between the center point trajectory divergence rate
 174 and the estimate from our 3D flights appears to be due to the distance from
 175 the point of estimation, rather than inconsistencies in the flights path due to
 176 buffeting. This can also be seen in table 3.1, where the correlation coefficients
 177 between the simulated flight and the 2D circle are all > 0.95 , while we see
 178 a steady drop in the correlation coefficients with the center point trajectory
 179 divergence rate as the radius increases.

180 Figure 4 shows the results for the attraction rate. Using data from the
 181 850mb isosurface we calculated the attraction rate at the center point of our

	2km circle	5km circle	10km circle	15km circle	2km flight	5km flight	10km flight	15km flight
center point	0.955	0.854	0.790	0.730	0.931	0.827	0.781	0.730
2km circle	--	0.946	0.815	0.751	0.981	0.923	0.811	0.765
5km circle		--	0.866	0.768	0.935	0.981	0.865	0.784
10km circle			--	0.928	0.804	0.836	0.974	0.902
15km circle				--	0.745	0.738	0.904	0.955
2km flight					--	0.944	0.824	0.783
5km flight						--	0.870	0.793
10km flight							--	0.937

Table 1: Pearson correlation coefficients for $\dot{\rho}$ measurements. Coefficients range from 0.730 to 0.965.

	2km circle	5km circle	10km circle	15km circle	2km flight	5km flight	10km flight	15km flight
center point	0.939	0.838	0.677	0.590	0.910	0.821	0.675	0.577
2km circle	--	0.932	0.742	0.644	0.980	0.917	0.739	0.627
5km circle		--	0.898	0.789	0.916	0.980	0.887	0.760
10km circle			--	0.908	0.729	0.881	0.978	0.864
15km circle				--	0.637	0.788	0.907	0.965
2km flight					--	0.936	0.746	0.644
5km flight						--	0.900	0.791
10km flight							--	0.909

Table 2: Pearson correlation coefficients for attraction rate measurements. Coefficients range from 0.577 to 0.939.

182 circle/flight radius, shown in red. We then used velocity data from a perfectly
183 circular path with a radius varying from 2km to 15km restricted to the 850mb
184 isosurface to approximate the attraction rate, shown in black. Finally, we
185 used velocity data from a 3D simulated UAS flight path with a radius varying
186 from 2km to 15km attempting to follow the 850mb isosurface to approximate
187 the attraction rate, shown in blue. Pearson correlation coefficients for these
188 measurements can be found in table 3.1.

189 We can see from the results in figure 4 that the simulated UAS flight
190 in a 3D space provides a very similar attraction rate measurements to the
191 circular path restricted to the 850mb isosurface. For all the radii paths we
192 looked at the attraction rate from the flight simulation is nearly identical to
193 that from the 2D circular path. Most of the error between the center point
194 attraction rate and the estimate from our 3D flights is due to the distance
195 from the point of estimation, rather than inconsistencies in the flights path
196 due to buffeting. This can also be seen in table 3.1, where the correlation
197 coefficients between the simulated flight and the 2D circle are all > 0.96 ,
198 while we see a steep drop in the correlation coefficients with the center point
199 attraction rate as the radius increases.

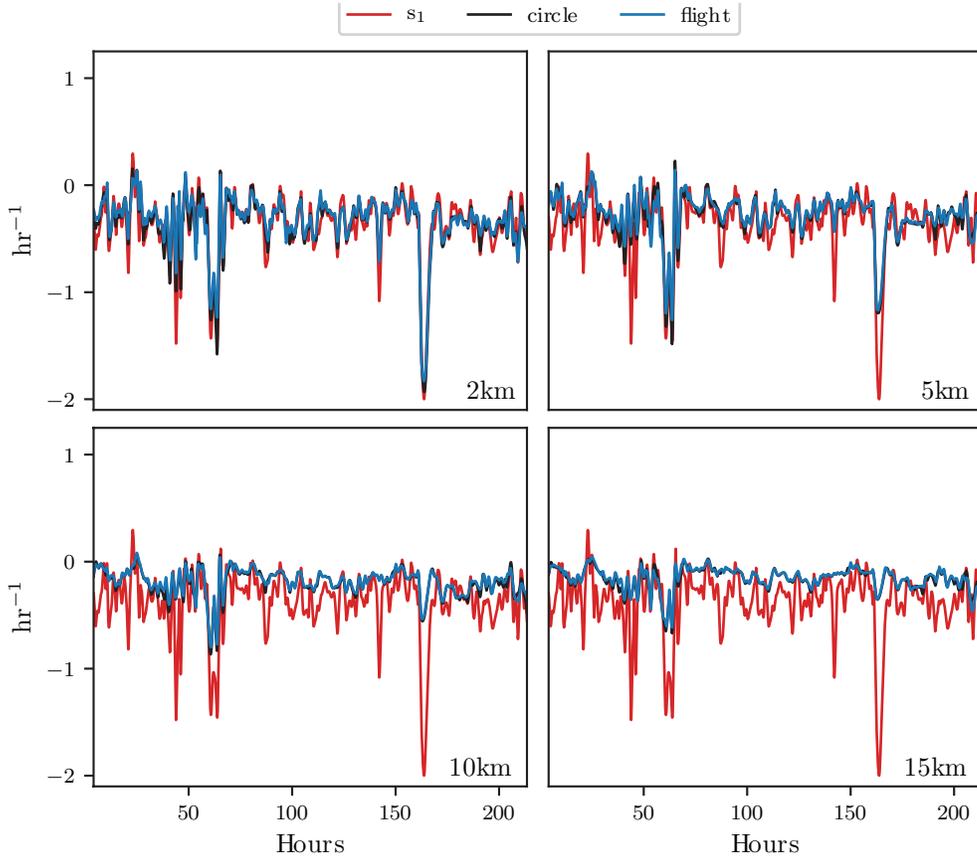


Figure 4: Comparison of the attraction rate measurements between center point (red), circular arc (black), and simulated drone flight (blue). Radius for the measurements is shown in lower right hand corner.

200 Both the attraction rate and the trajectory divergence rate at a point can
 201 be approximated to a high degree of accuracy by UAS flights. Simulated 3D
 202 UAS flights provided measurements which were nearly identical to those of
 203 perfect circular 2D paths. The main cause of error in the approximations
 204 appears to be the distance of the path from the center point. Furthermore,
 205 the trajectory divergence rate appears to be a more robust metric than the
 206 attraction rate; meaning that the trajectory divergence rate can be better
 207 approximated at larger radii than the attraction rate can. This can be seen
 208 very clearly seen in tables 1 and 2, where the correlation coefficient for the
 209 attraction rate drops off much quicker than for the trajectory divergence rate.

210 **3.2 Using Eulerian Metrics to infer Lagrangian Dy-** 211 **namics**

212 In this section we examine how well the attraction rate, s_1 , and the trajectory
213 divergence rate, $\dot{\rho}$, do at predicting Lagrangian dynamics, such as the passage
214 of LCSs. Figure 5 shows the time series for the trajectory divergence rate
215 and backward-time FTLE for integration times of 0.5, 1, and 2 hrs. The
216 FTLE values have been multiplied by -1 for improved visualization. In this
217 figure we can see that the trajectory divergence rate does not always follow
218 the trend of the negative backward-time FTLE, which is to be expected. The
219 trajectory divergence rate gives information on both instantaneous attraction
220 and repulsion, while the negative backward-time FTLE gives a measure of
221 attraction. The trajectory divergence rate does, however, agree with the
222 negative backward-time FTLE when we have significant periods of attraction.
223 This behavior is of particular interest for the detection of LCSs. When
224 calculating LCSs, there is often a multitude of weaker, less important LCSs.
225 In order to filter out these less important structures and focus on important
226 structures, one often needs to set a threshold value for the FTLE field. These
227 dips in the the trajectory divergence rate, coinciding with the strongest dips
228 in the negative backward-time FTLE, would therefore seem to be a likely
229 indicator of LCS of interest.

230 Figure 6 shows the time series for the attraction rate and backward-
231 time FTLE for integration times of 0.5, 1, and 2 hrs. The FTLE values
232 have been multiplied by -1 for improved visualization. In this figure we
233 can see that the attraction rate follows the general trend of the negative
234 backward-time FTLE. This makes sense as both the attraction rate and the
235 negative backward-time FTLE give measures of attraction. The attraction
236 rate therefore, should give a good approximation to the negative backward-
237 time FTLE, and thus should be able to give indications of LCSs.

238 We can further explore the effectiveness of the attraction rate and the
239 trajectory divergence rate for detecting LCSs by looking at receiver operating
240 characteristic (ROC) curves. For this we looked at when LCSs passed within
241 a threshold radius which ranged from 400m to 10km of our center point,
242 figure 7. We further applied a threshold of 90% for the LCSs, so only LCSs
243 whose FTLE value was within the 90th percentile were considered. We looked
244 at the attraction rate's and the trajectory divergence rate's ability to detect
245 LCSs for integration times of $\frac{1}{2}$, 1, and 2 hrs in backward-time.

246 Figure 8 show ROC curves for the the trajectory divergence rate. The

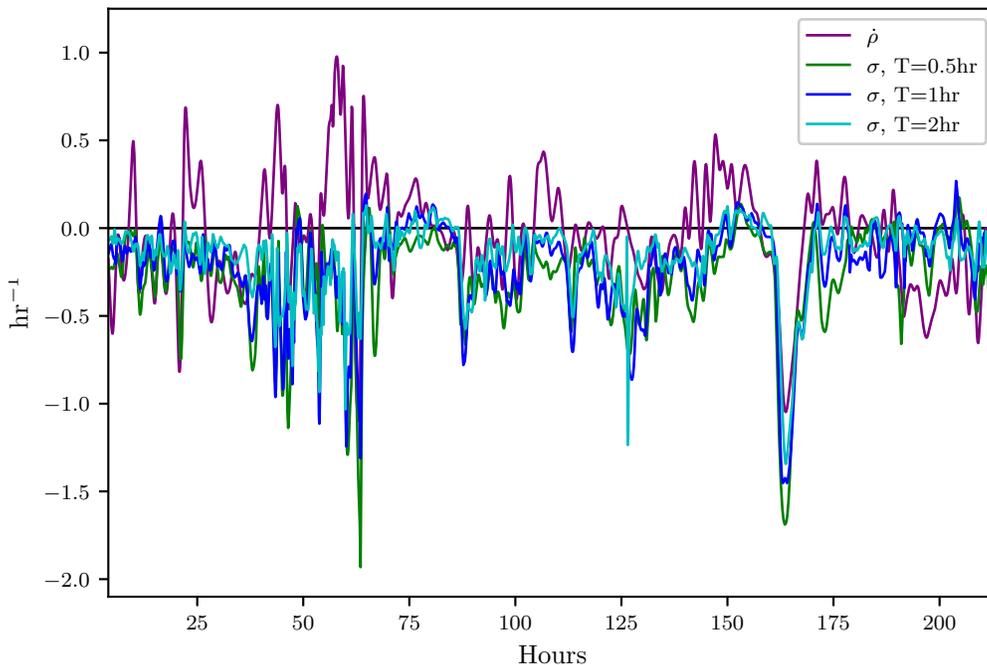


Figure 5: Comparison of the trajectory divergence rate with the 0.5, 1, and 2 hr backward-time FTLE from $t=4$ to $t=215$ hrs. FTLE fields have been multiplied by -1 to offer better comparison of attraction.

247 trajectory divergence rate gives measures of both attraction and repulsion,
 248 so to filter out repulsive indicators we first masked trajectory divergence rate
 249 values > 0 . After this, we threshold the from 0%, upper right hand side, to
 250 100%, lower left hand corner. Every 20th percentile is marked with a dot.
 251 Each subplot represents a different threshold radius, with radii ranging from
 252 400m to 10km. Each color represents a different integration time for the
 253 LCSs, 0.5hr green, 1hr red, 2hr blue. These ROC curves indicate that the
 254 trajectory divergence rate can indeed be used to detect LCSs passing through
 255 an area.

256 Figure 9 shows ROC curves for the attraction rate. We threshold the
 257 attraction rate from 0%, upper right hand corner, to 100%, lower left hand
 258 corner. Every 20th percentile is marked with a dot. Each subplot represents
 259 a different threshold radius, with radii ranging from 400m to 10km. Each
 260 color represents a different integration time for the LCSs, 0.5hr green, 1hr
 261 red, 2hr blue. These ROC curves indicate that not only can the attraction

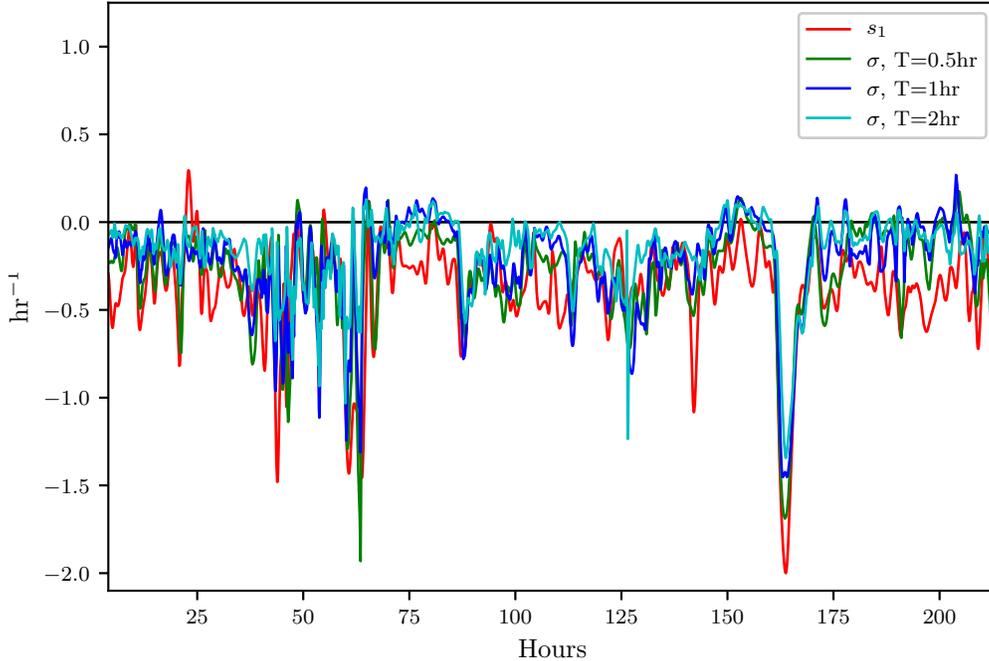


Figure 6: Comparison of the attraction rate with the 0.5, 1, and 2 hr backward-time FTLE from $t=4$ to $t=215$ hrs. FTLE fields have been multiplied by -1 to offer better comparison of attraction.

262 rate be used to detect LCSs passing through an area, but it does a better
 263 job detecting attracting LCSs than the trajectory divergence rate does.

264 It should be noted that both the attraction and trajectory divergence
 265 rates seem to perform best at an area threshold of around 800-2000m and
 266 converge to random chance as the radius increases. We suspect that this
 267 is due to the spatial and temporal scales of the input data, 3km x 1hr grid
 268 spacing. We speculate that with a velocity field continuously defined in space
 269 and time, we would see continued improvement in the ROC curves as the
 270 threshold radius decreases. Unfortunately the analytic models currently used
 271 in the study of LCSs, such as the double gyre [29] and the Bickley jet [30], do
 272 not have the requisite spatial in-homogeneity necessary to reveal meaningful
 273 Eulerian structures.

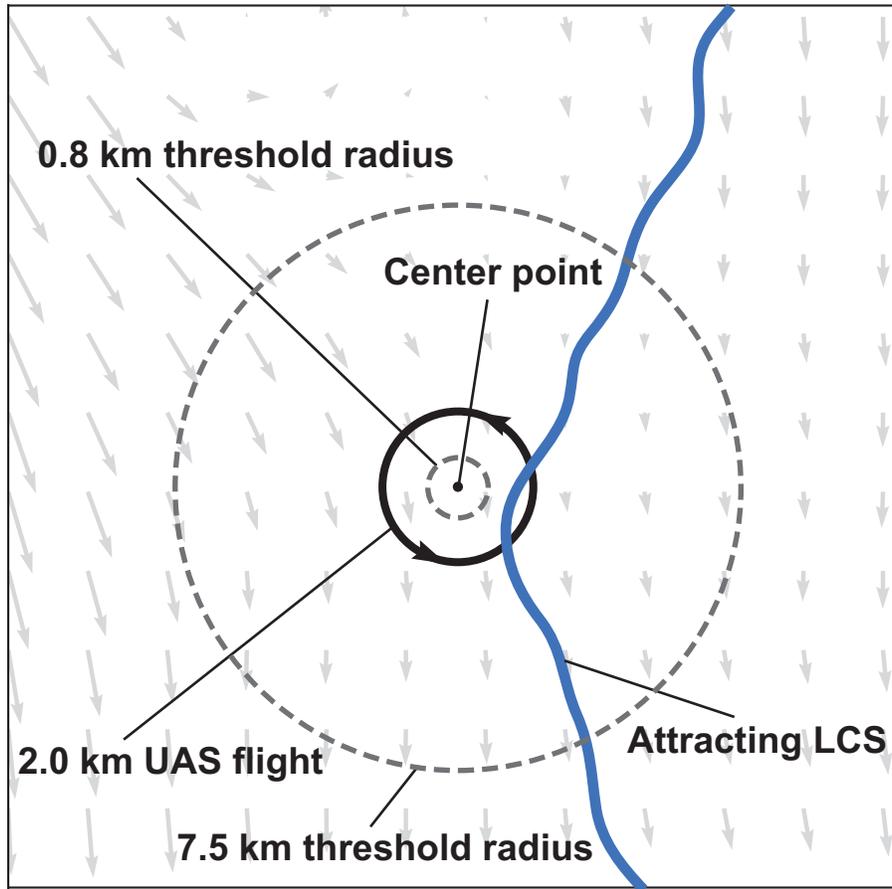


Figure 7: schematic of LCS detection. Two examples of threshold radii shown as dashed lines, with an attracting LCS falling within one of threshold radius yet outside the other.

274 **3.3 Inferring Lagrangian Dynamics from UAS mea-**
 275 **surements**

276 In this section we examine how well the attraction rate, s_1 , and the trajec-
 277 tory divergence rate, $\dot{\rho}$ as approximated from a UAS flight do at predicting
 278 Lagrangian dynamics, such as the passage of LCSs. Figure 10 shows ROC
 279 curves for the trajectory divergence rate as calculated from a simulated
 280 2km UAS flight. Once again we first masked trajectory divergence rate val-
 281 ues > 0 to filter out repulsive indicators. After this, we threshold the from

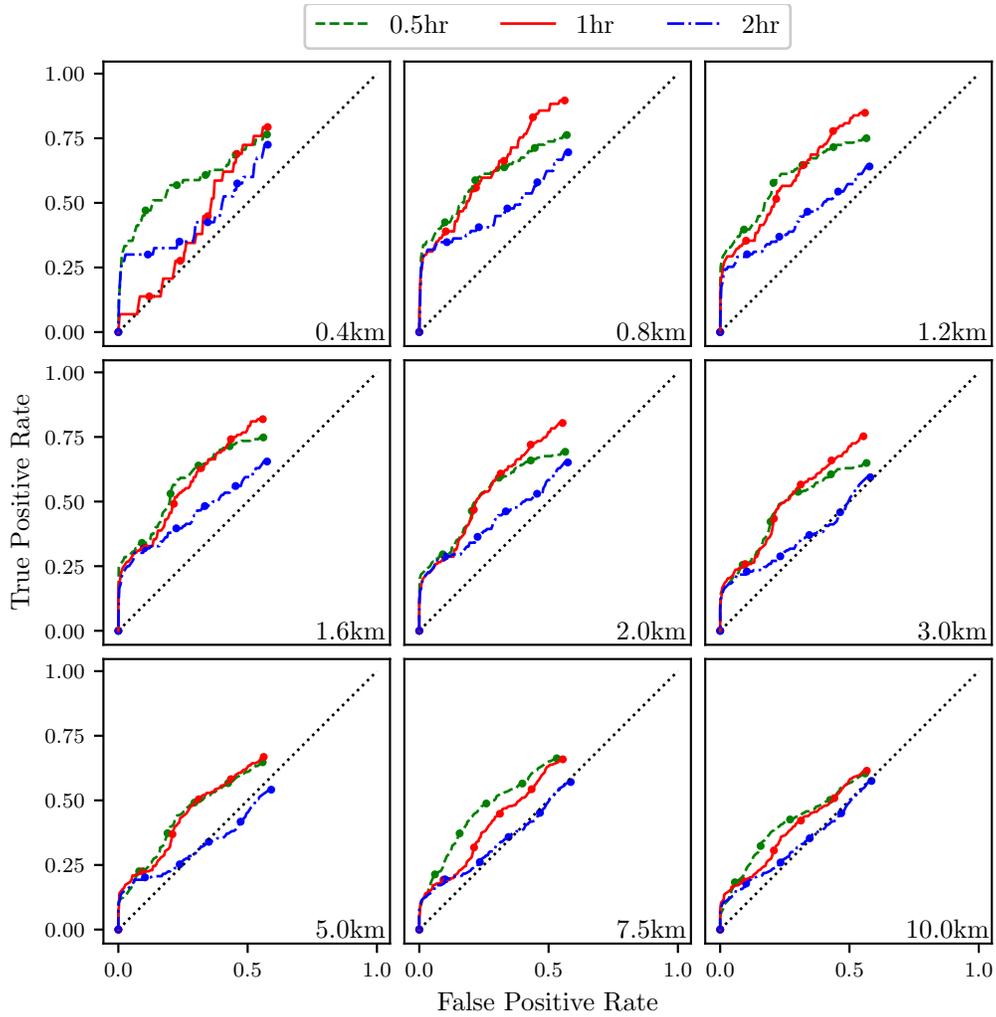


Figure 8: ROC curves for the trajectory divergence rate as measured at the center point ability to detect 90% percentile LCSs with integration times of 0.5 (green), 1 (red), and 2 (blue) hrs. Threshold radii are displayed in the lower left hand corner.

282 0%, upper right hand side, to 100%, lower left hand corner. Every 20th per-
 283 centile is marked with a dot. Each subplot represents a different threshold
 284 radius, with radii ranging from 400m to 10km. Each color represents a dif-
 285 ferent integration time for the LCSs, 0.5hr green, 1hr red, 2hr blue. These
 286 ROC curves show a striking resemblance to the ROC curves in figure 8. This

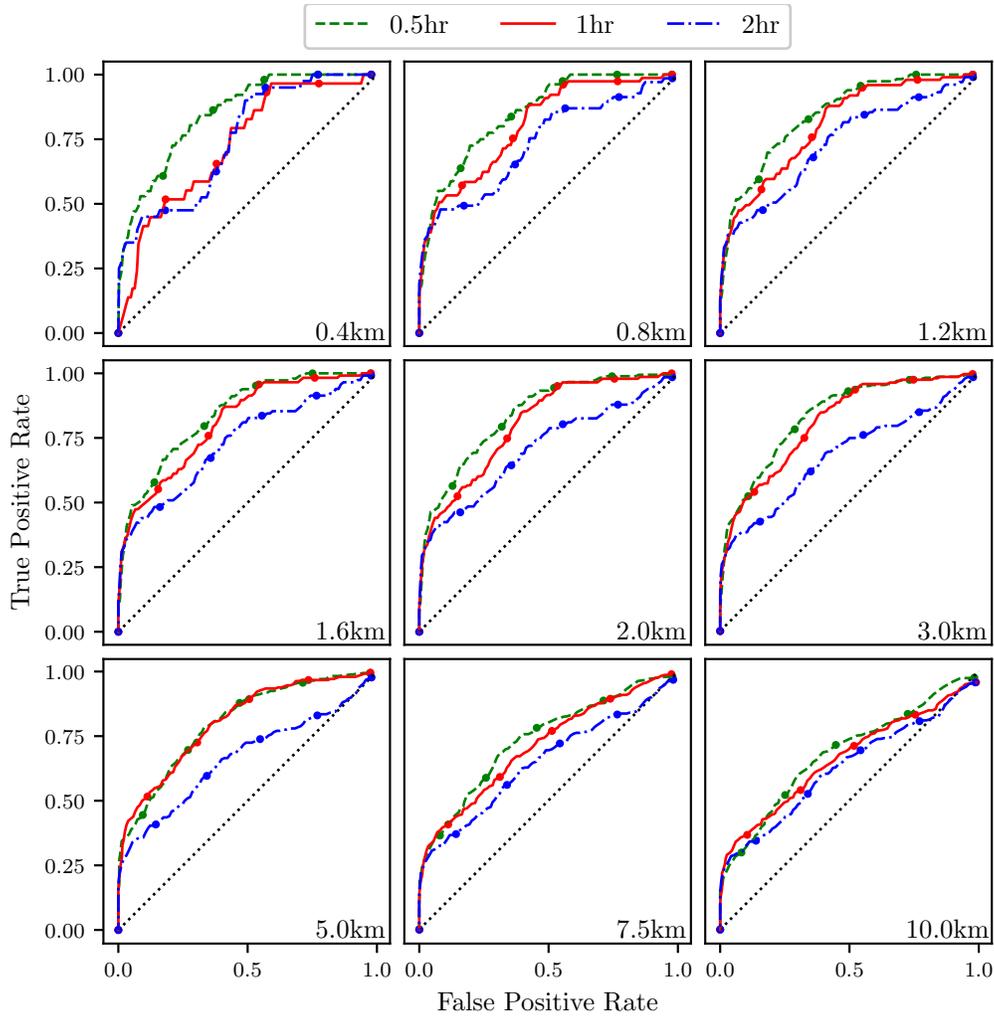


Figure 9: ROC curves for the attraction rate as measured at the center point ability to detect 90% percentile LCSs with integration times of 0.5 (green), 1 (red), and 2 (blue) hrs. Threshold radii are displayed in the lower left hand corner.

287 would indicate that the trajectory divergence rate as approximated from a
 288 UAS flight can indeed be used to detect LCSs passing through an area.

289 Figure 11 shows ROC curves for the attraction rate as calculated from
 290 a simulated 2km UAS flight. We threshold the attraction rate field from
 291 0%, upper right hand corner, to 100%, lower left hand corner. Every 20th

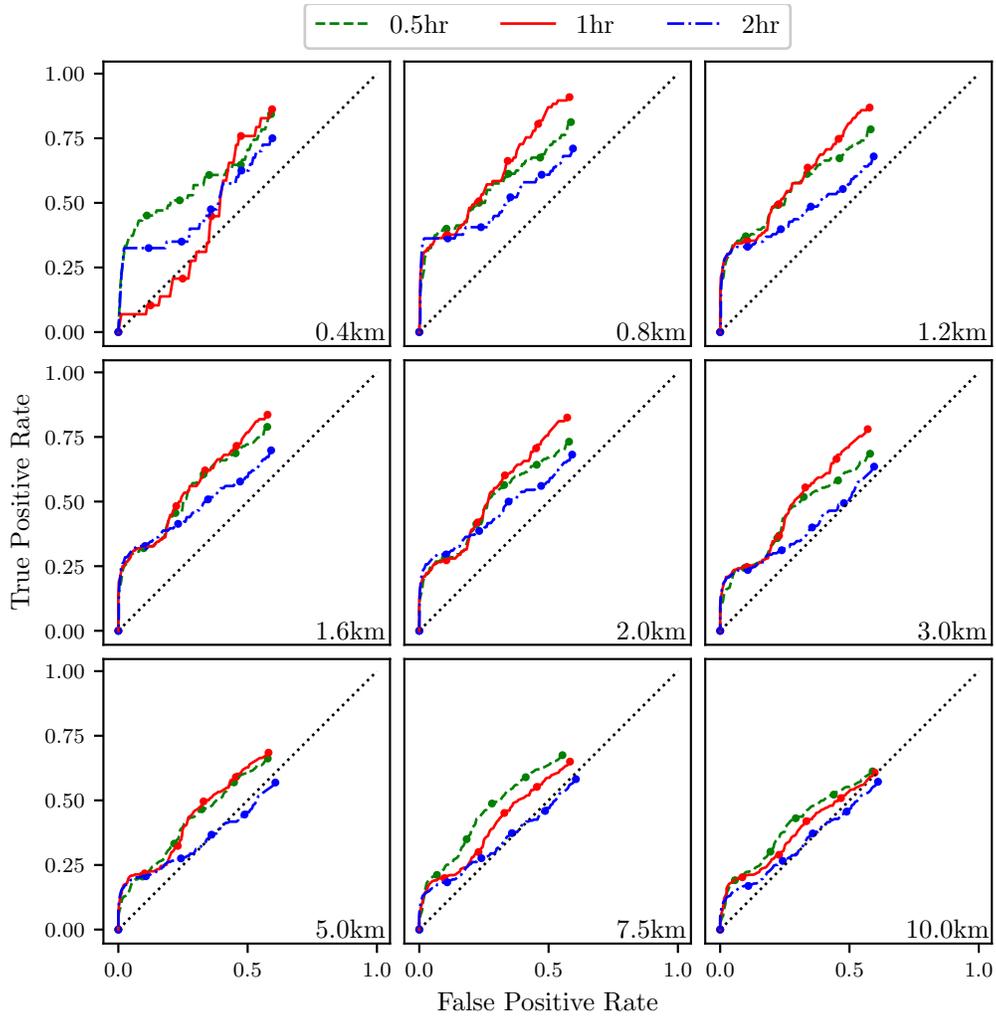


Figure 10: ROC curves for the trajectory divergence rate as measured from a 2km radius UAS simulation ability to detect 90% percentile LCSs with integration times of 0.5 (green), 1 (red), and 2 (blue) hrs. Threshold radii are displayed in the lower left hand corner.

292 percentile is marked with a dot. Each subplot represents a different thresh-
 293 old radius, with radii ranging from 400m to 10km. Each color represents a
 294 different integration time for the LCSs, 0.5hr green, 1hr red, 2hr blue. As
 295 before, these ROC curves closely resemble the ROC curves in figure 9. This
 296 would indicate that the attraction rate as approximated from a UAS flight

297 can also be used to detect LCSs passing through an area.

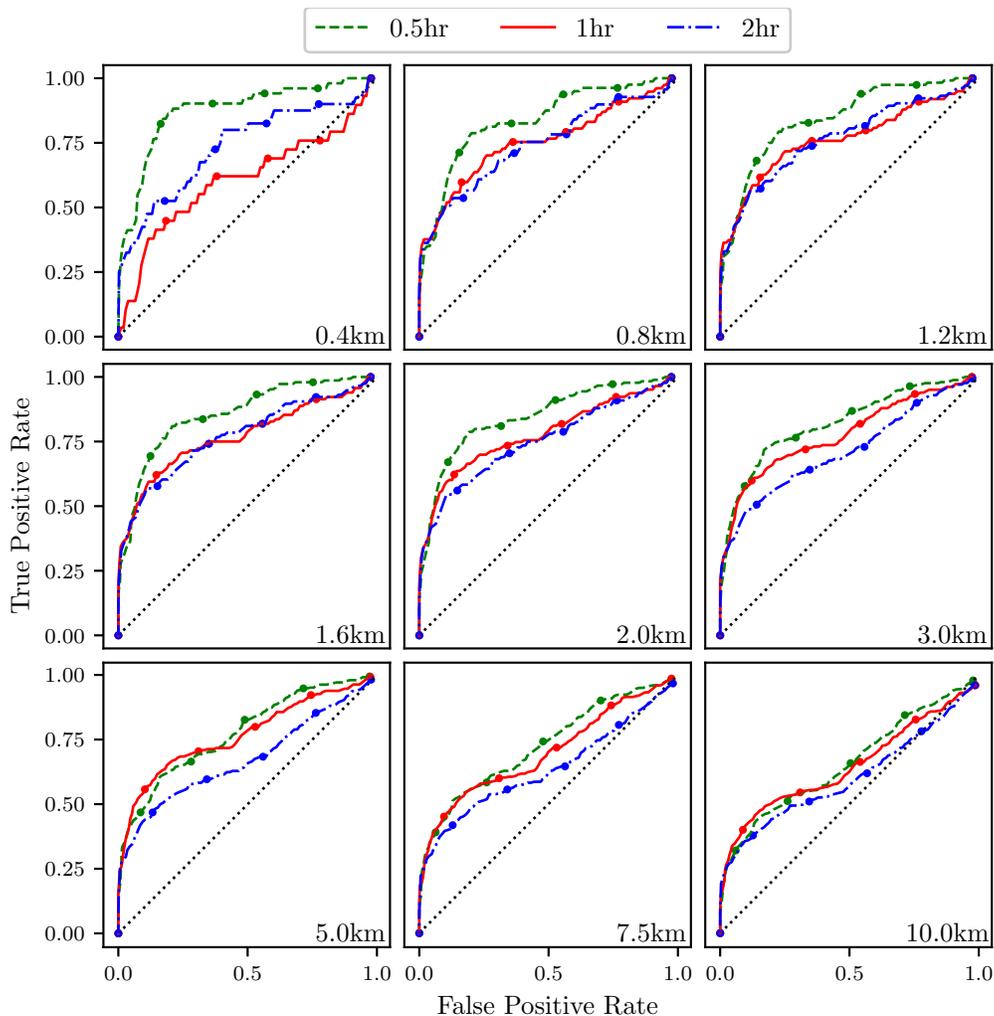


Figure 11: ROC curves for the attraction rate as measured from a 2km radius UAS simulation ability to detect 90% percentile LCSs with integration times of 0.5 (green), 1 (red), and 2 (blue) hrs. Threshold radii are displayed in the lower left hand corner.

298 The ROC curves in figures 10 and 11 indicate that both the trajectory
299 divergence rate and the attraction rate as approximated from a UAS flight
300 can be used to infer local Lagrangian dynamics. Furthermore, the attraction
301 rate appears to be a much better indicator of passing attractive LCSs than

302 the trajectory divergence rate. Interestingly, at around a 70-90% threshold
303 the attraction rate as approximated by the UAS flight, figure 11, seems to
304 outperforms the attraction rate at the center point, figure 9. We suspect
305 that this is due to the fact that the UAS measurements are taking in infor-
306 mation from a larger area and with these higher thresholds is filtering out
307 the additional noise.

308 4 Conclusion

309 We have put forward a novel algorithm to approximate the gradient of a
310 scalar field using measurements from a circular arc around a point. Using
311 realistic atmospheric velocity data from the NAM 3km model, we applied
312 this algorithm to circular trajectories restricted a 2D isosurface and simu-
313 lated UAS flights in 3D, with radii ranging from 2km to 15km. From these
314 results we approximated the trajectory divergence rate and the attraction
315 rate for the center point of these paths. Comparing these approximations
316 with the trajectory divergence rate and attraction rate at the center point,
317 we found that both the flight and the circle gave nearly identical approxi-
318 mations. Furthermore, the approximations were very good for the smaller
319 radii we looked, but even the larger radii approximations were able to pick
320 up the trend of the trajectory divergence rate and attraction rate, though
321 they underestimated the magnitude.

322 We have also examined the ability of Eulerian diagnostics, in particular
323 the trajectory divergence and attraction rates, to infer Lagrangian dynamics.
324 Using ROC curves, we first looked at the ability of the trajectory divergence
325 rate and attraction rate, as measured at a point to detect the passage of
326 LCSs within a threshold radius. We found that the attraction rate can be
327 used as an effective tool to sense short term LCS passing by. We also found
328 that the trajectory divergence rate, while performing better than chance,
329 under performed the attraction rate. We then extended this to look at the
330 trajectory divergence rate and attraction rate as approximated by a UAS
331 flight. Once again we found that these Eulerian diagnostics, as approximated
332 by a UAS flight, can be an effective tool for detecting LCSs passing through
333 a sampling area.

334 This paper serves as a first step in real-time detection of LCSs in the
335 atmosphere. It demonstrates that a fixed wing UAS can, in principle, be
336 used to measure Eulerian diagnostics of a local atmospheric flow. These

337 Eulerian diagnostics can then be used to infer the Lagrangian dynamics of
 338 the local flow. Future work will apply this to real world data to detect actual
 339 atmospheric LCSs, evaluate the effects of sensor uncertainty on the accuracy
 340 of LCS detection, and extend this to the detection of pollutant specific LCSs,
 341 such as those found along atmospheric rivers [12].

342 A Flight Dynamic Model

343 The aircraft flight dynamic model comes from combining standard aircraft
 344 rigid-body equations [31] with Grauer and Morelli’s Generic Global Aero-
 345 dynamic model [32], modified for non-uniform wind. The important flight
 346 dynamic modeling assumptions are:

- 347 1. Earth is a flat, inertial reference.
- 348 2. The aircraft is a rigid body, symmetric about its longitudinal plane,
 349 with constant mass m .
- 350 3. For wind-aircraft interaction, the aircraft is a point “located” at it’s
 351 center-of-mass.
- 352 4. The wind is described by a C^1 -smooth kinematic vector field.
- 353 5. Aircraft thrust T is an instantaneously-controllable force acting nose-
 354 forward from the center-of-mass.
- 355 6. All parameters are invariant with altitude. (e.g. no altitudinal variation
 356 of density ρ , gravity g , ground-effect, etc.)

357 The resulting dynamic equations of motion are

$$\mathbf{R}_{BM}(\alpha) \begin{pmatrix} C_D(\dots) \\ C_Y(\dots) \\ C_L(\dots) \end{pmatrix} \frac{1}{2} \rho \|\mathbf{V}_r\|^2 S + \begin{pmatrix} T\% \left(\frac{T_{max}}{100\%} \right) \\ 0 \\ 0 \end{pmatrix} + \mathbf{R}_{BE}(\Theta) \begin{pmatrix} 0 \\ 0 \\ mg \end{pmatrix} = m \left(\dot{\mathbf{V}} + (\boldsymbol{\omega} \times \mathbf{V}) \right), \quad (8)$$

$$\begin{bmatrix} b & 0 & 0 \\ 0 & \bar{c} & 0 \\ 0 & 0 & b \end{bmatrix} \begin{pmatrix} C_l(\dots) \\ C_m(\dots) \\ C_n(\dots) \end{pmatrix} \frac{1}{2} \rho \|\mathbf{V}_r\|^2 S = \begin{bmatrix} I_{xx} & 0 & -I_{xz} \\ 0 & I_{yy} & 0 \\ -I_{xz} & 0 & I_{zz} \end{bmatrix} \dot{\boldsymbol{\omega}} + \left(\boldsymbol{\omega} \times \begin{bmatrix} I_{xx} & 0 & -I_{xz} \\ 0 & I_{yy} & 0 \\ -I_{xz} & 0 & I_{zz} \end{bmatrix} \boldsymbol{\omega} \right). \quad (9)$$

358 where the ellipses on the aerodynamic coefficients remind the reader that
 359 these are functions of state variables, as given below in Equations 12 – 17.
 360 The symbol V is used for inertially-referenced velocity, and V_r is used for air-
 361 relative velocity. These dynamics are combined with standard translational
 362 and rotational kinematic equations

$$\dot{X} = \mathbf{R}_{EB}(\phi, \theta, \psi) \begin{pmatrix} u \\ v \\ w \end{pmatrix} = \mathbf{R}_{EB}(\Theta) \mathbf{V}, \quad (10)$$

363

$$\dot{\Theta} = \begin{pmatrix} \dot{\phi} \\ \dot{\theta} \\ \dot{\psi} \end{pmatrix} = \begin{bmatrix} 1 & \sin \phi \tan \theta & \cos \phi \tan \theta \\ 0 & \cos \phi & -\sin \phi \\ 0 & \sin \phi \sec \theta & \cos \phi \sec \theta \end{bmatrix} \begin{pmatrix} p \\ q \\ r \end{pmatrix} = \mathbf{L}^{-1}(\phi, \theta) \boldsymbol{\omega}. \quad (11)$$

364 The aerodynamic coefficient expressions are from Equation 20 of Grauer and
 365 Morelli [32]

$$C_D = \theta_1 + \theta_2 \alpha + \theta_3 \alpha \tilde{q}_r + \theta_4 \alpha \delta_e + \theta_5 \alpha^2 + \theta_6 \alpha^2 \tilde{q}_r + \theta_7 \alpha^2 \delta_e + \theta_8 \alpha^3 + \theta_9 \alpha^3 \tilde{q}_r + \theta_{10} \alpha^4, \quad (12)$$

$$C_Y = \theta_{11} \beta + \theta_{12} \tilde{p}_r + \theta_{13} \tilde{r}_r + \theta_{14} \delta_a + \theta_{14} \delta_r, \quad (13)$$

$$C_L = \theta_{16} + \theta_{17} \alpha + \theta_{18} \tilde{q}_r + \theta_{19} \delta_e + \theta_{20} \alpha \tilde{q}_r + \theta_{21} \alpha^2 + \theta_{22} \alpha^3 + \theta_{23} \alpha^4, \quad (14)$$

$$C_l = \theta_{24} \beta + \theta_{25} \tilde{p}_r + \theta_{26} \tilde{r}_r + \theta_{27} \delta_a + \theta_{28} \delta_r, \quad (15)$$

$$C_m = \theta_{29} + \theta_{30} \alpha + \theta_{31} \tilde{q}_r + \theta_{32} \delta_e + \theta_{33} \alpha \tilde{q}_r + \theta_{34} \alpha^2 \tilde{q}_r + \theta_{35} \alpha^2 \delta_e + \theta_{36} \alpha^3 \tilde{q}_r + \theta_{37} \alpha^3 \delta_e + \theta_{38} \alpha^4, \quad (16)$$

$$C_n = \theta_{39} \beta + \theta_{40} \tilde{p}_r + \theta_{41} \tilde{r}_r + \theta_{42} \delta_a + \theta_{43} \delta_r + \theta_{44} \beta^2 + \theta_{45} \beta^3. \quad (17)$$

366 In these equations $(\theta_1, \theta_2, \dots, \theta_{45})$ are the aircraft parameters, (α, β) are the
 367 standard aerodynamic angles, $(\tilde{p}_r, \tilde{q}_r, \tilde{r}_r)$ are wind-relative non-dimensionalized
 368 angular rates, and $(\delta_a, \delta_e, \delta_r)$ are aileron, elevator, and rudder deflections.

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