

# Evaluation of the Numerical Values of the Normalization Parameter $a$ and the Distance Scale $\alpha$ .

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Here we will determine the numerical values of  $a$  and  $\alpha$  in order to determine with more precisions the impact of dark energy in the vicinity of a black hole.

We know that the metric of a Schwarzschild black in the quintessence field is given by

$$ds^2 = -g(r)dt^2 + \frac{1}{g(r)}dr^2 + r^2(d\theta^2 + \sin^2\theta d\varphi^2),$$

with

$$g(r) = 1 - \frac{2GM}{c^2} \frac{1}{r} - \frac{a}{r^{3\epsilon+1}}.$$

Then, for the reason of homogeneities, we can think that  $a$  is proportional to a power of radius. So, at the event horizon, we have the normalization factor that way

$$a \propto r_h^{3\epsilon+1}.$$

Then, if the proportional factor is 1, we could have

$$a = r_h^{3\epsilon+1} = \left(\frac{2GM}{c^2}\right)^{3\epsilon+1}.$$

More over we have written for the distance scale  $\alpha$  that

$$\alpha \propto a^{\frac{1}{3\epsilon+1}}.$$

Now, we have to determine the true range of  $a$  and  $\alpha$ . For that, we know that

$$\begin{aligned} -1 &< \epsilon \leq -0.30 \\ G &= 6.67408 * 10^{-11} m^3.kg^{-1}s^{-2} \\ C &= 299792458 m.s^{-1} \end{aligned}$$

For a stellar black hole with the mass  $M = 4M_{sun} = 7.9564 * 10^{30} kg$ , we have this board

$\epsilon$	-0.99	-0.80	-0.70	-0.50	-0.40	-0.30
$a$	$9.4881 * 10^{-9}$	$1.9884 * 10^{-6}$	$3.3132 * 10^{-5}$	$9.1992 * 10^{-3}$	$1.5328 * 10^{-1}$	2.5542
$\alpha$ (metre)	11816.7442	11816.7302	11816.8612	11816.7998	11818.7771	11818.1366

For this board we see that for any value of  $\epsilon$  in the range  $]-1, -0.30]$ , we have some specifics values of the normalization parameter  $a$  and the distance scale  $\alpha$ , permitting us to appreciate better the influence of the dark energy in the vicinity of black holes.

### Bibliography

**Ndongmo et al, Black Hole Mass Decreasing, The power and The Time of Two Black Holes in Coalescence *viXra:1811.0097***