## Using metaballs to model the merger of Schwarzschild black holes

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## Abstract

In this paper, Blinn's metaballs are used to model the merger of Schwarzschild black holes.

## 1 Metaballs

Metaballs have been used in computer graphics ever since their discovery by Jim Blinn.

Where  $G = c = \hbar = k = 1$ , there is an analytical solution for the merger of n metaballs (Schwarzschild black holes) travelling directly toward each other at non-relativistic speeds:

$$f(l) = \sum_{i=1}^{n} \frac{2M_i}{r_i},\tag{1}$$

where  $M_i$  is the mass of the *i*th metaball, and

$$r_i = \sqrt{(l.x - v_i.x)^2 + (l.y - v_i.y)^2 + (l.z - v_i.z)^2},$$
(2)

where l is the sample location, and  $v_i$  is the centre of the ith metaball. The isosurface (event horizon) is given by

$$f(l) = 1. (3)$$

Included are figures of a black hole merger. The isosurface was tessellated using the Marching Cubes algorithm.

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Figure 1: Two black holes of unit mass each, at a distance of 11.



Figure 2: Two black holes of unit mass each, at a distance of 10.

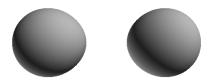


Figure 3: Two black holes of unit mass each, at a distance of 9.



Figure 4: Two black holes of unit mass each, at a distance of 8.



Figure 5: Two black holes of unit mass each, at a distance of 7.



Figure 6: Two black holes of unit mass each, at a distance of 6.



Figure 7: Two black holes of unit mass each, at a distance of 5.



Figure 8: Two black holes of unit mass each, at a distance of 4.



Figure 9: Two black holes of unit mass each, at a distance of 3.



Figure 10: Two black holes of unit mass each, at a distance of 2.



Figure 11: Two black holes of unit mass each, at a distance of 1.



Figure 12: One black hole of mass = 2.

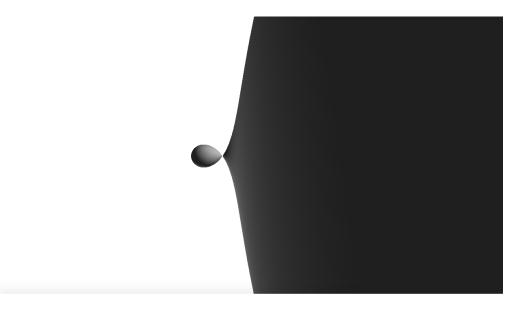


Figure 13: A unit mass black hole merging with a black hole of mass = 1000.

## References

- [1] Emparan R, Martinez M. Exact event horizon of a black hole merger (2016) arXiv:1603.00712 [gr-qc]
- [2] Schutz B. A First Course in General Relativity (Cambridge: Cambridge University Press, 1985)