Cosmological constant of GRT as a radial function in dependence of velocity

- A short notice -

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Abstract:

Under special circumstances cosmological "constant" of GRT can be formulated as a function in dependence of radial term. This calculation will be shown. In fact this system of physical ideas is now described only for local state of Schwarzschild-lineelement with cosmological variable but it can be easily developed to cosmic terms.

Key-words: Cosmological constant; Einstein-equation; gravity-equation; Schwarzschild-solution, flat space-time; Planck-length; radial function.

<u>1. Introduction:</u>

Since Einstein introduced this cosmological term Λ to correct and complete his gravity equations in 1917 ad hoc for logically consistent description in four spacetime dimensions [1.], this term plays an important role in description of universe in its global states, particularly as a form of "dark energy", which determines the observed acceleration of cosmic expansion or can be interpreted as a form of a vacuum-energy. Mostly this term is considered as a constant but, as is shown, it also can be interpreted as a function in dependence of radius.

2. Calculation:

If the Schwarzschild-lineelement of a local spacetime is written with cosmological constant [2.]:

$$ds^{2} = \frac{dr^{2}}{1 - \frac{2 \cdot M}{r} - \frac{\Lambda \cdot r^{2}}{3}} + r^{2} \cdot (d\theta^{2} + \sin^{2}(\theta) \cdot d\phi^{2}) - c^{2} \cdot dt^{2} \left(1 - \frac{2M}{r} - \frac{\Lambda \cdot r^{2}}{3}\right)$$
(1.)

where: $M = \frac{2 \cdot G \cdot m}{c^2}$ is Schwarzschild-radius with *m* central-matter-mass of gravity-field

which causes the material part of the gravity-field [3.] and the limit is now done for

 $m=0 \Rightarrow M=0$, then the lineelement is describing a local flat form of spacetime without a centralmass but with the cosmological term Λ . From materia its empty like a geon, first formulated by Wheeler [4.]. This g- field now can be described far from its empty source by setting:

 $ds^2 = dr_{PL}^2$ (far out in the wilderness) as its physical minimal size.

This field then can be written as:

$$dr_{PL}^{2} = \frac{dr^{2}}{\left(1 - \frac{\Lambda \cdot r^{2}}{3}\right)} - c^{2} \cdot dt^{2} \cdot \left(1 - \frac{\Lambda \cdot r^{2}}{3}\right)$$
(2.)

which leads directly to:

$$\Lambda = \frac{1}{r^2} + \frac{3}{2 \cdot r^2} \cdot \frac{dr_{PL}^2}{2 \cdot c^2 \cdot dt^2} \pm \frac{3}{r^2} \sqrt{\frac{dr_{PL}^4}{4 \cdot c^4 \cdot dt^4} + \frac{v^2}{c^2}}$$
(3.)

where in local spacetime is defined:

 $\frac{dr^2}{c^2 \cdot dt^2} := \frac{v^2}{c^2}.$ neglecting the term *d* in space-and timelike coordinate-differentials because it can be left out of consideration for this theme.

<u>3. Conclusion:</u>

Cosmological term Λ can be written as a function, which depends local on the variables of velocity v and radial variable r resp. timelike differential dt. In "classical" GRT without Planck-length as a fundamental minimal length with the continuity condition $\hbar \Rightarrow 0$, this term reduces then to:

$$\Lambda = \frac{1}{r^2} \cdot \left(1 \pm \frac{3 \cdot v}{c} \right) \tag{4.}$$

If this function can be also interpreted as a global cosmic description, then dark energy can't be a constant but must depend from cosmical expansion-radius and in interpretation from cosmical expansion-velocity.

Solution:
$$\Lambda = \Lambda(r, v) \neq const$$
. (5.)

4. Summary:

The cosmological term Λ of GRT can be written as a function in dependence from velocity and of radius. This result comes from explanation of a local examination in Schwarzschild-lineelement with cosmological-term but can be developed in an explanation to global cosmic expansion like is actually written and observed in [5.].

5. Comment:

Since Ricci-scalar is coupled with cosmological term via

$$\Lambda = \frac{\chi \cdot T - R}{4} \tag{6.}$$

where χ is Einstein-gravitational constant and T is $diag \sum_{i=1}^{4} T_{i}^{k}$; i = k, Ricci-scalar then

can also be written as a function from distance r and velocity v. With the assumption $m=0 \Rightarrow M=0$ there is also T=0. This leads to a result for Ricci-scalar as a function of R(r,v,dt):

$$R = \frac{-4}{r^2} - \frac{3}{r^2} \cdot \frac{dr_{PL}^2}{c^2 \cdot dt^2} \pm \frac{12}{r^2} \cdot \sqrt{\frac{dr_{PL}^4}{4 \cdot c^4 \cdot dt^4}} + \frac{v^2}{c^2} \qquad .$$
(7.)

which reduces in classical GRT with the continuity condition $\hbar \Rightarrow 0$ to:

$$R = \frac{-4}{r^2} \cdot \left(1 \pm \frac{3 \cdot v}{c} \right) \quad . \tag{8.}$$

5. References:

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6.Verification:

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